
What long-term data can tell us about Molniya: past works and open points

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MOLNIYA CONSTELLATION

VERY RICH DYNAMICAL ENVIRONMENT:

a perfect laboratory for different dynamical systems tools and perspectives

- **Semi-synchronous orbit:** 12 hour orbital period — tesseral resonance
- **Highly elliptical orbit:** $a \approx 26550$ km, $e \approx 0.73$ — lunisolar perturbation and atmospheric drag
- **Critical inclination:** $i \approx 63.4$ degrees — $\dot{\omega}_{J_2} \approx 0$



LITERATURE

- L. Anselmo, C. Pardini, Long-term simulation of objects in high-earth orbits, ESA/ESOC Study Note (2006).
- P. Christopher, Molniya system alternatives for geostationary satellite systems with applications to 72-100GHz systems, in Proc. Ka Broadband Conf., Cagliari, Italy, 2009
- D.G. King-Hele, The orbital lifetimes of Molniya satellites, Journal of the British Interplanetary Society 28 (1975) 783-796
- Y.F. Kolyuka, N.M. Ivanov, T.I. Afanasieva, T.A. Gridchina, Examination of the lifetime, evolution and re-entry features for the "Molniya" type orbits, International Symposium on Space Flight Dynamics, Toulouse, France, September 28-October 2, 2009
- E. Kuznetsov, P. Zakharova, Dynamical evolution of space debris on high-elliptical orbits near high-order resonance zones, Advances in Space Research 56 (2015) 406-413
- M.L. Lidov, A.A. Solovov, Some qualitative regularities and evaluations of orbital evolution of Molniya 1 satellites, Cosmic Research 13 (1976) 709-716
- J.T. McGraw, P.C. Zimmer, M.R. Ackermann, Ever Wonder What's in Molniya? We do., Advanced Maui Optical and Space Surveillance (AMOS) Technologies Conference, September 19-22, 2017, Maui, Hawaii
- J.P. Murphy, E.L. Victor, A determination of the second and fourth order sectorial harmonics in the geopotential from the motion of 12-hr satellites, Planetary and Space Science 16 (1968) 195-204
- A.I. Nazarenko, I.V. Usovik, Instability of the solution of the problem on determining the reentry time of satellites on elliptic orbits, Acta Astronautica 163 (2019) 142-146
- J. Silha, T. Schildknecht, A. Hinze, T. Flohrer, A. Vananti, An optical survey for space debris on highly eccentric and inclined MEO orbits 2017, Advances in Space Research 59 (2017) 181-192
- A.S. Sochilina, On the motion of a satellite in resonance with its rotating planet, Celestial Mechanics 26 (1982) 337-352
- R.-y. Sun, C.-y. Zhao, M.-j. Zhang, Y.-G. Hou, Dynamical evolution of high area-to-mass ratio objects in Molniya orbits, Advances in Space Research 51 (2013) 2136-2144
- A.P. Trishchenko, L. Garand, Spatial and Temporal Sampling of Polar Regions from Two-Satellite System on Molniya Orbit, Journal of Atmospheric and Oceanic Technology 977 (2011) 977-992
- Yu.P. Ulybyshev, Design of satellite constellations with continuous coverage on elliptic orbits of Molniya type, Cosmic Research 47 (2009) 310-321
- T.-L. Zhu, C.-Y. Zhao, M.-J. Zhang, Long term evolution of Molniya orbit under the effect of Earth's non-spherical gravitational perturbation, Advances in Space Research 54 (2014) 197-208
- T.-L. Zhu, C.-Y. Zhao, H.-B Wang., M.-J. Zhang, Analysis on the long term orbital evolution of Molniya satellites, Astrophysics and Space Science 357 (2015) 126

ACKNOWLEDGMENTS

- J. Daquin, E.M. Alessi, J. O' Leary, A. Lemaitre, A. Buzzoni. Dynamical properties of the Molniya satellite constellation: long-term evolution of the semi-major axis, *Nonlinear Dynamics* 105 (2021), 2081-2103.
- T. Talu, E.M. Alessi, G. Tommei. On the Dominant Lunisolar Perturbations for Long-Term Eccentricity Variation: The Case of Molniya Satellite Orbits, *Universe* 7 (2021), 482.
- E.M. Alessi, A. Buzzoni, J. Daquin, A. Carbognani, G. Tommei. Dynamical properties of the Molniya satellite constellation: long-term evolution of orbital eccentricity, *Acta Astronautica* 179 (2021), 659-669.
- A. Buzzoni, J. Guichard, E.M. Alessi, G. Altavilla, A. Figer, A. Carbognani, G. Tommei. Spectrophotometric and dynamical properties of the Soviet/Russian constellation of Molniya satellites, *Journal of Space Safety Engineering* 7 (2020), 255-261.

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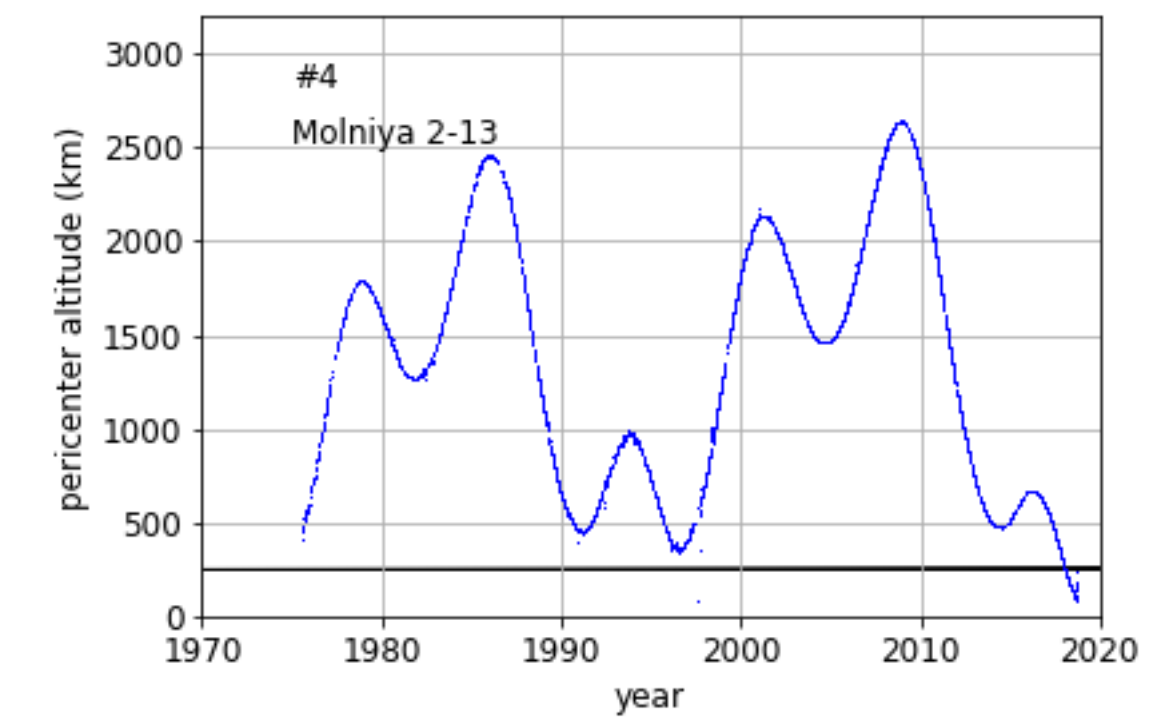
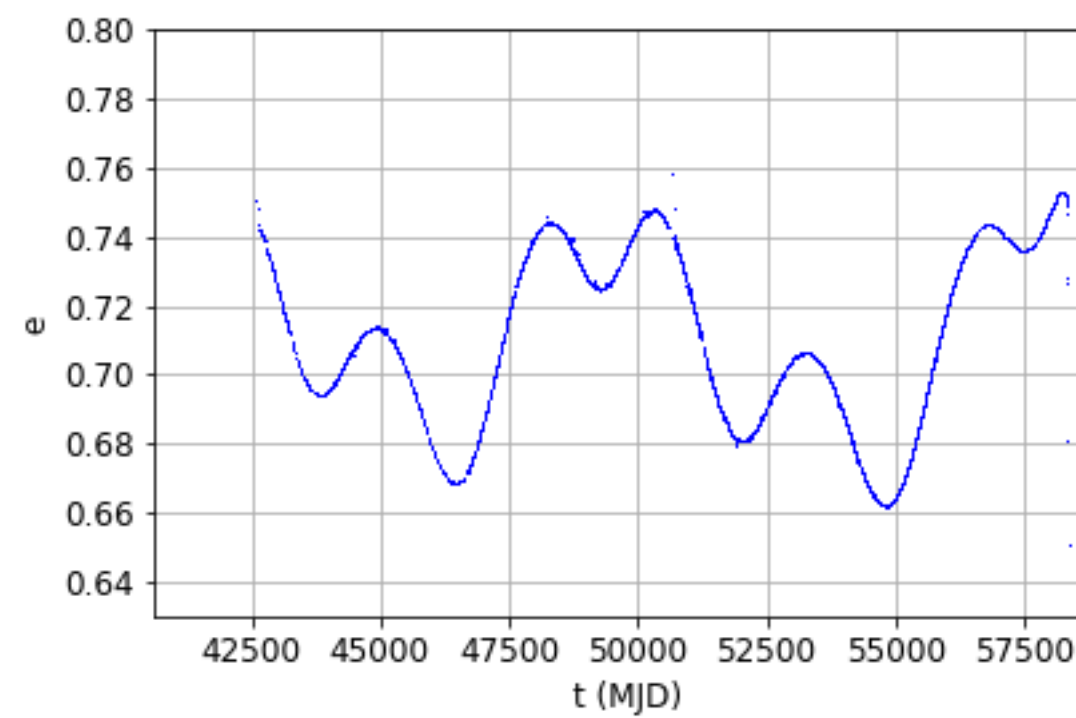
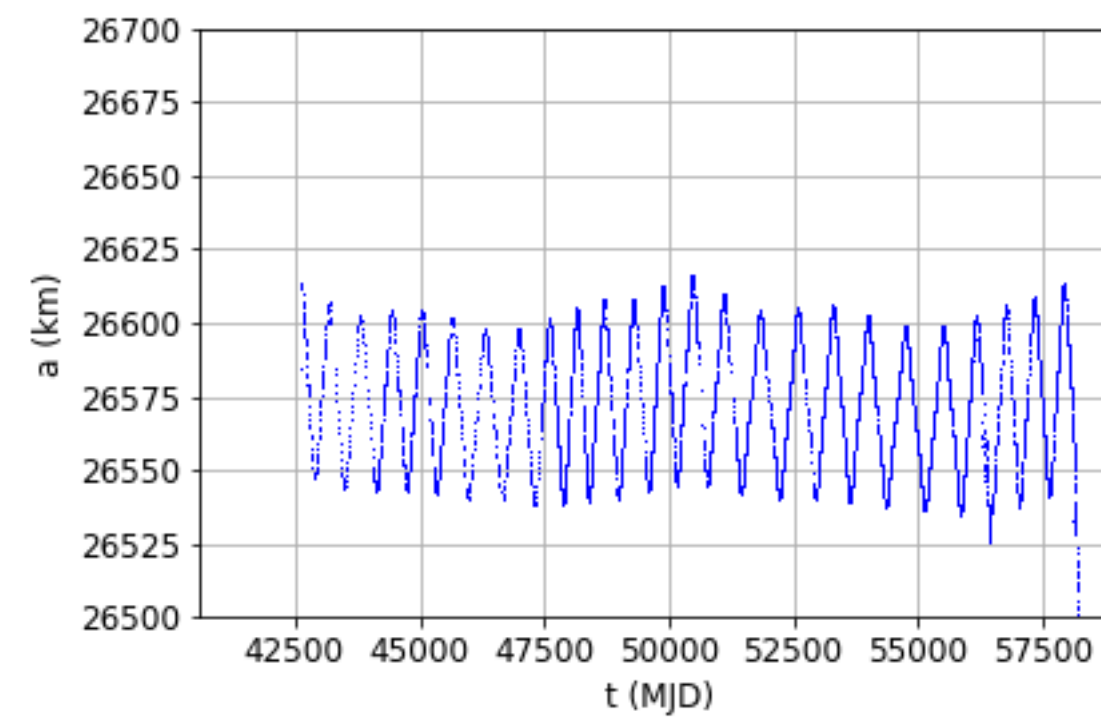
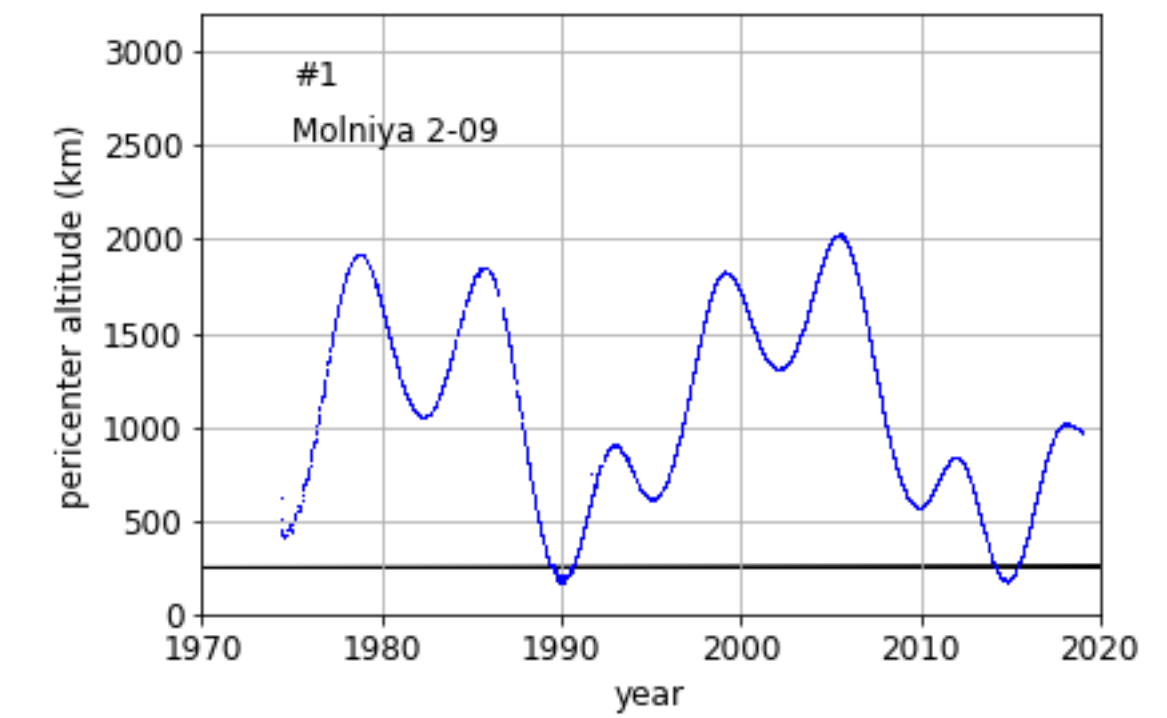
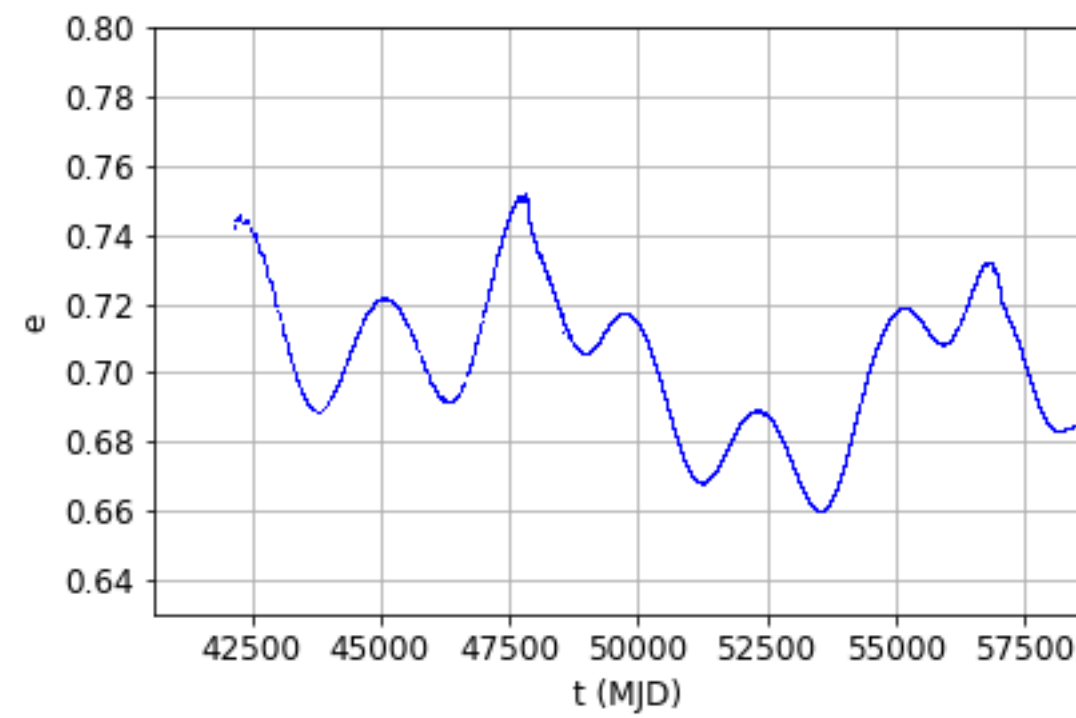
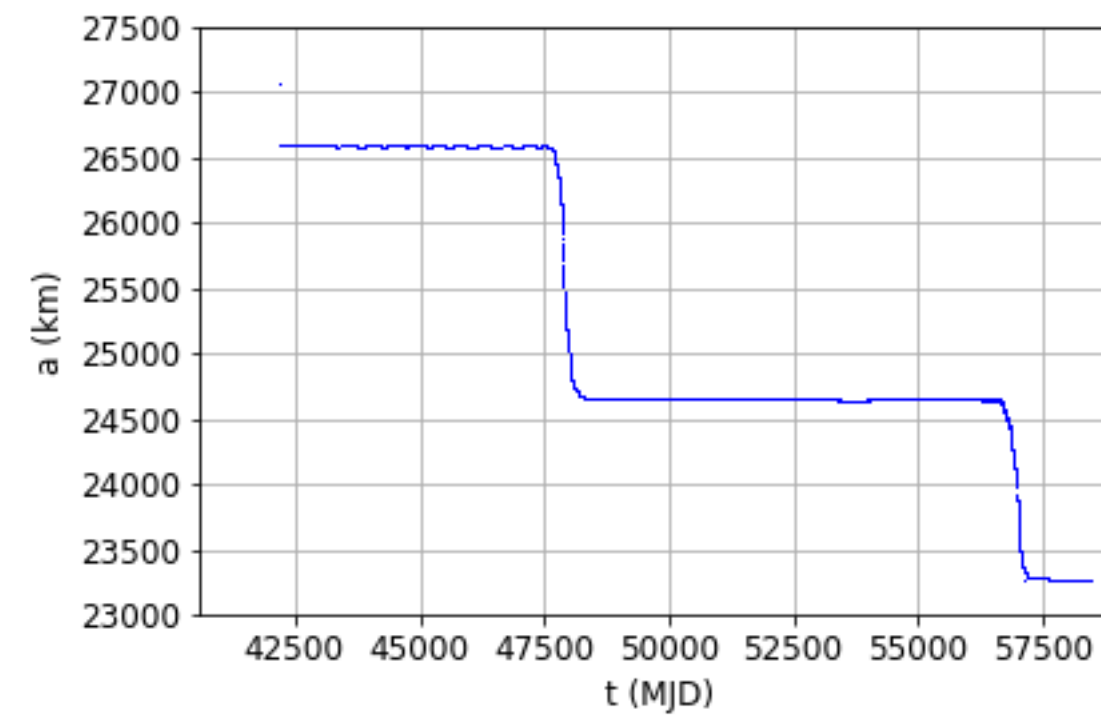
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AIM

The behavior in eccentricity and semi-major axis of 42 TLE sets was analyzed in order to:

- **get some of the main dynamical features of the orbits in the long term**
- **test different tools:**
 - hierarchy of dynamical models
 - numerical propagation
 - phase space portraits
 - frequency analysis
- **looking for a synergy between the pseudo-observational data and a natural perturbation approach**

EVOLUTION GIVEN BY THE TLEs

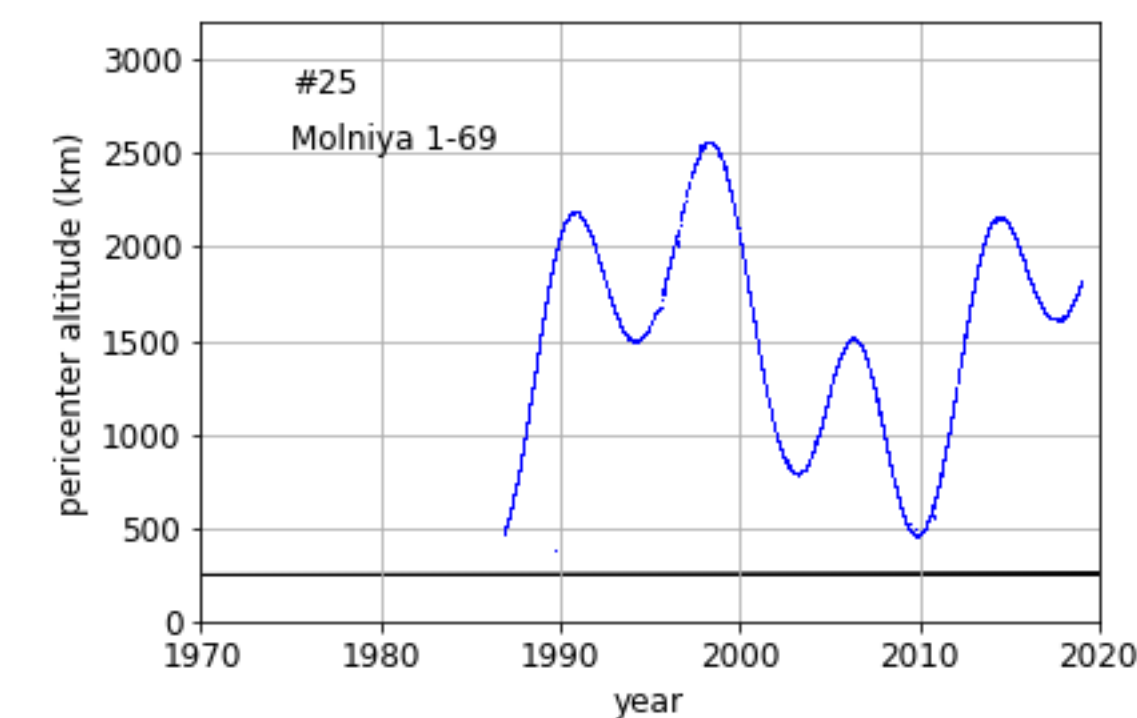
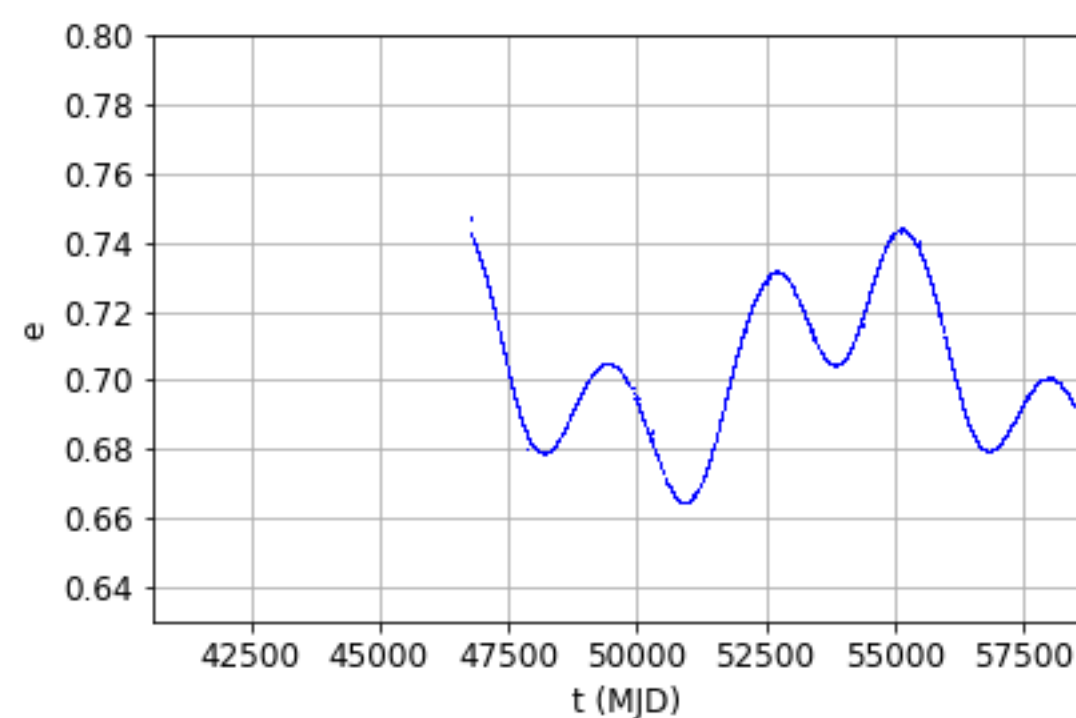
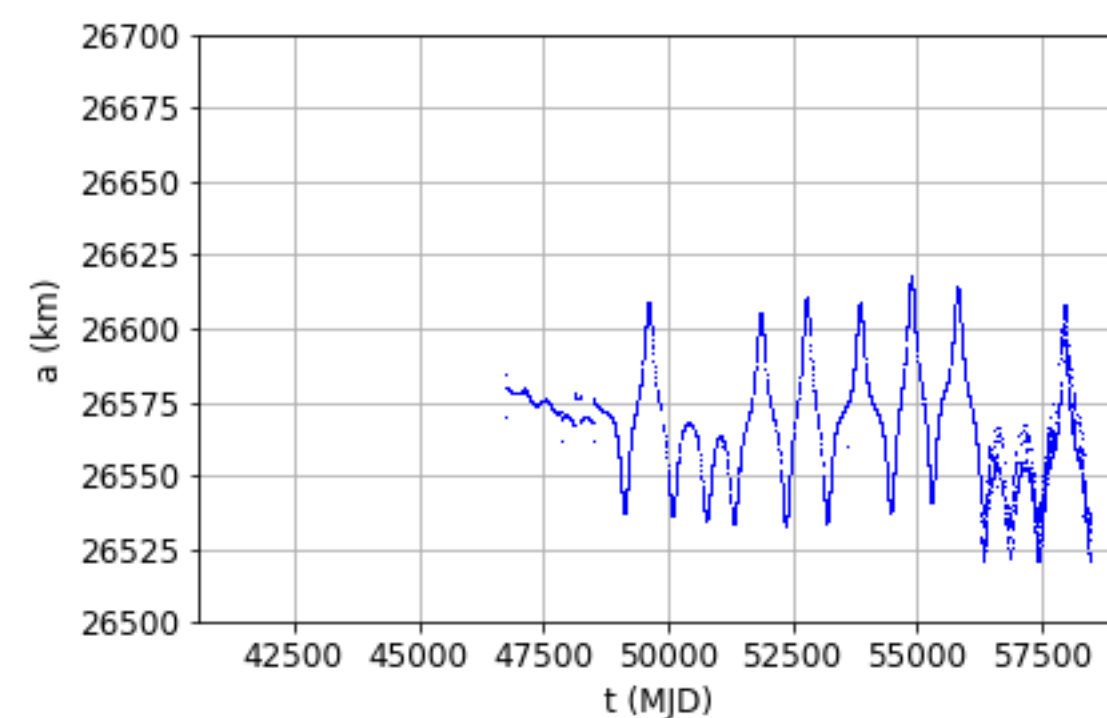
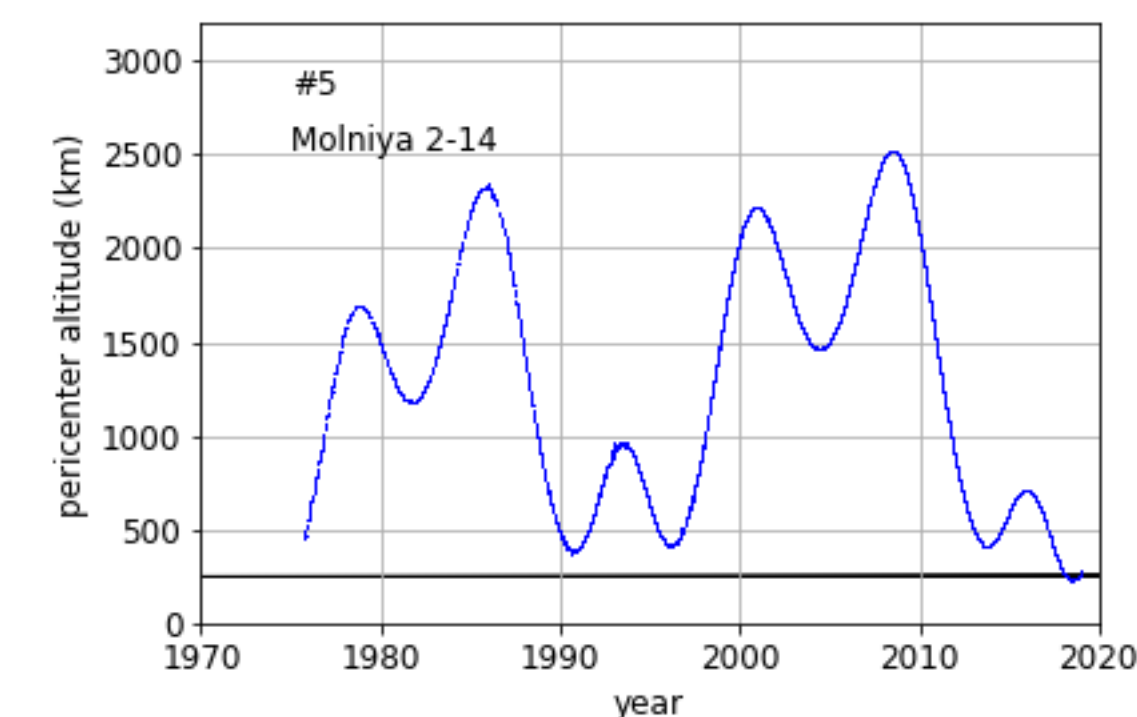
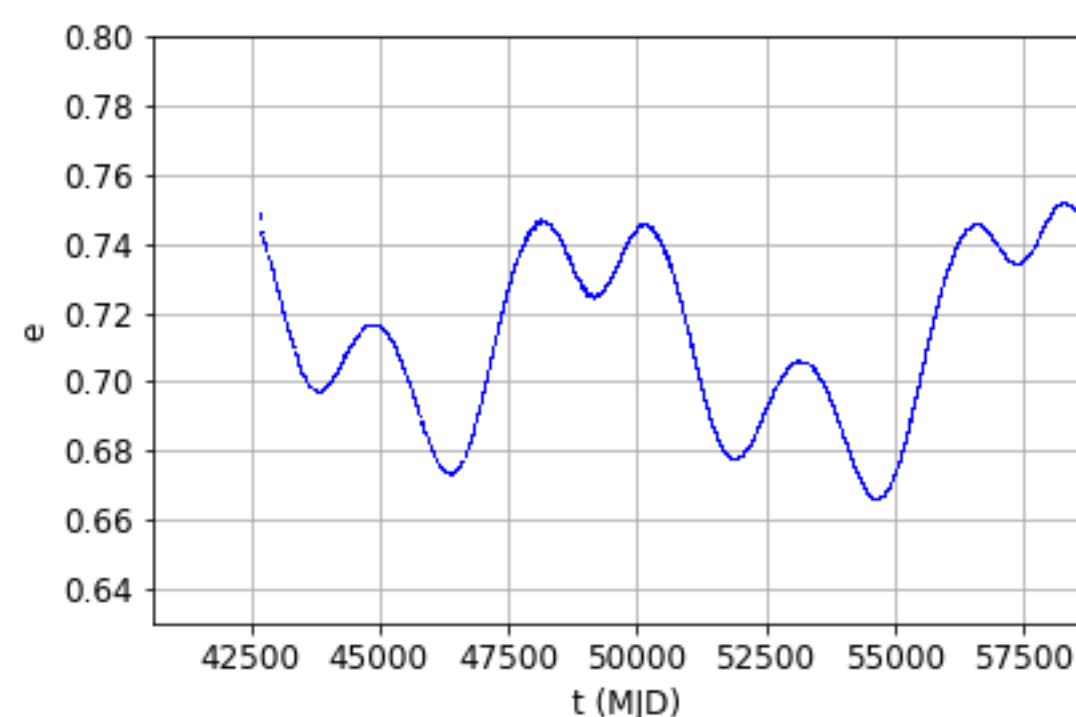
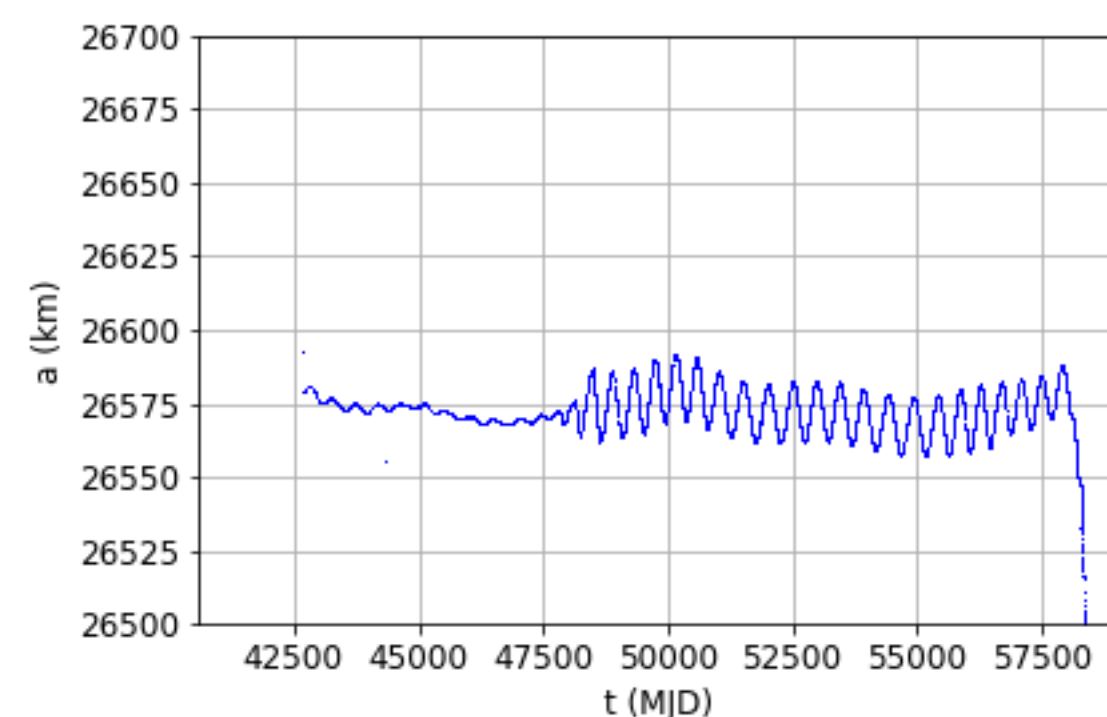


semi-major axis

eccentricity

pericenter altitude

EVOLUTION GIVEN BY THE TLEs



semi-major axis

eccentricity

pericenter altitude

ECCENTRICITY EVOLUTION

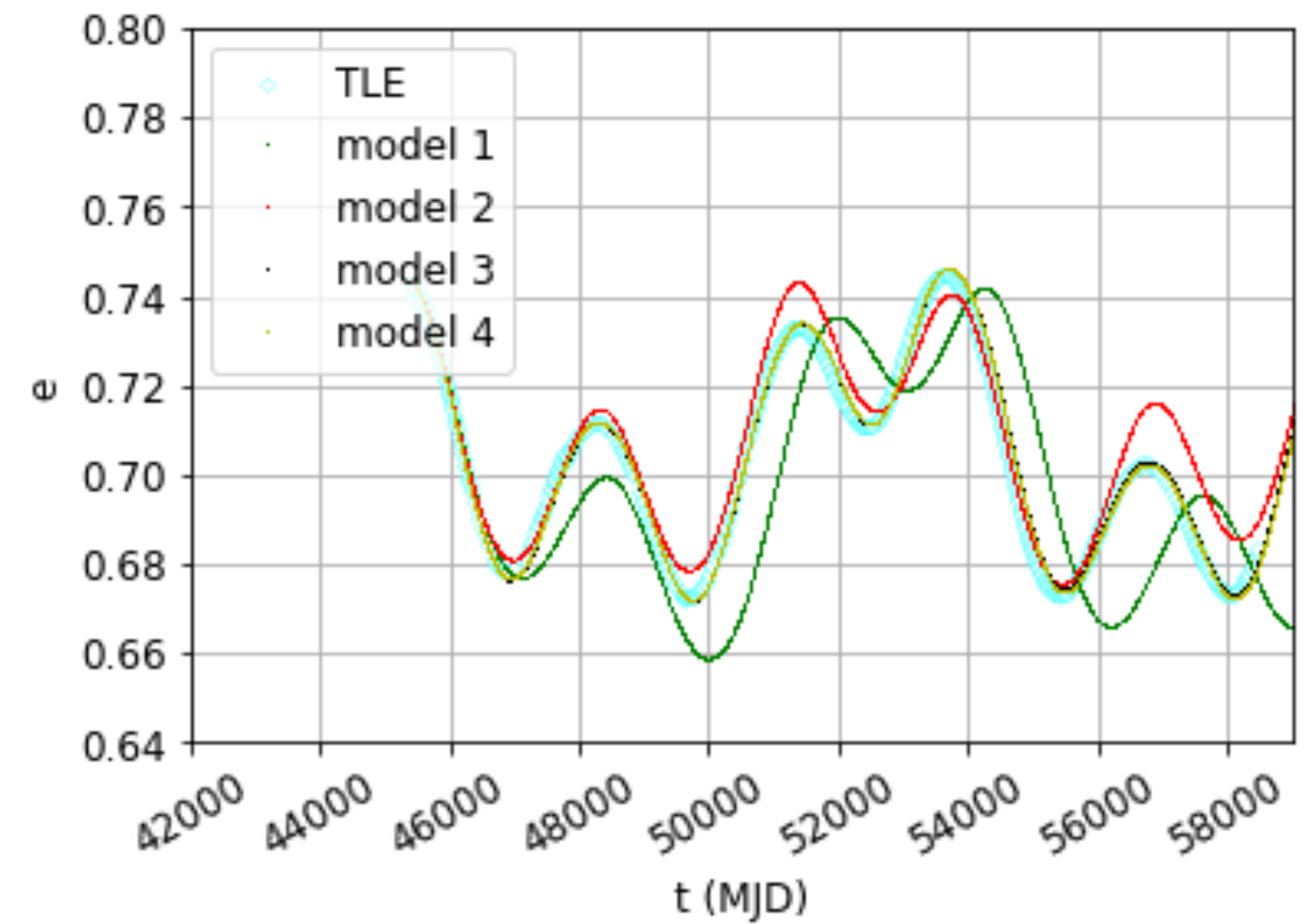
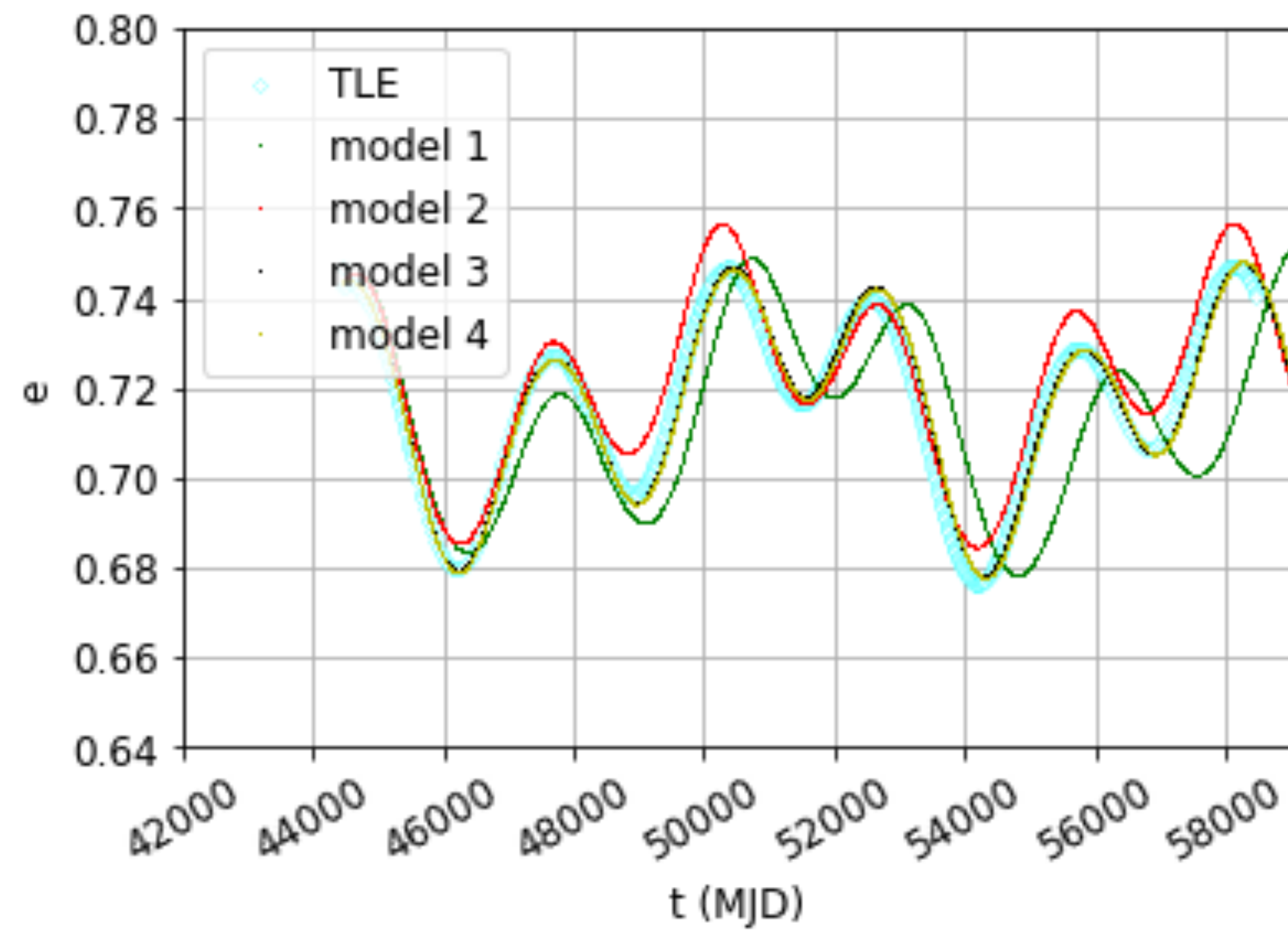
ECCENTRICITY EVOLUTION

FIRST APPROACH TO TACKLE THE LONG-TERM BEHAVIOR IN ECCENTRICITY:

*numerical propagation considering different models accounting only
for the oblateness effect and the lunisolar perturbation*

- **Doubly-averaged** approximation for the lunisolar perturbation
- Main assumption: **semi-major axis is constant**
- **Initial condition from the TLE, s.t. both $\dot{\omega}_{J_2} \approx 0$ and the 2:1 tesseral resonance are satisfied**
- Hierarchy of models

RESULTS



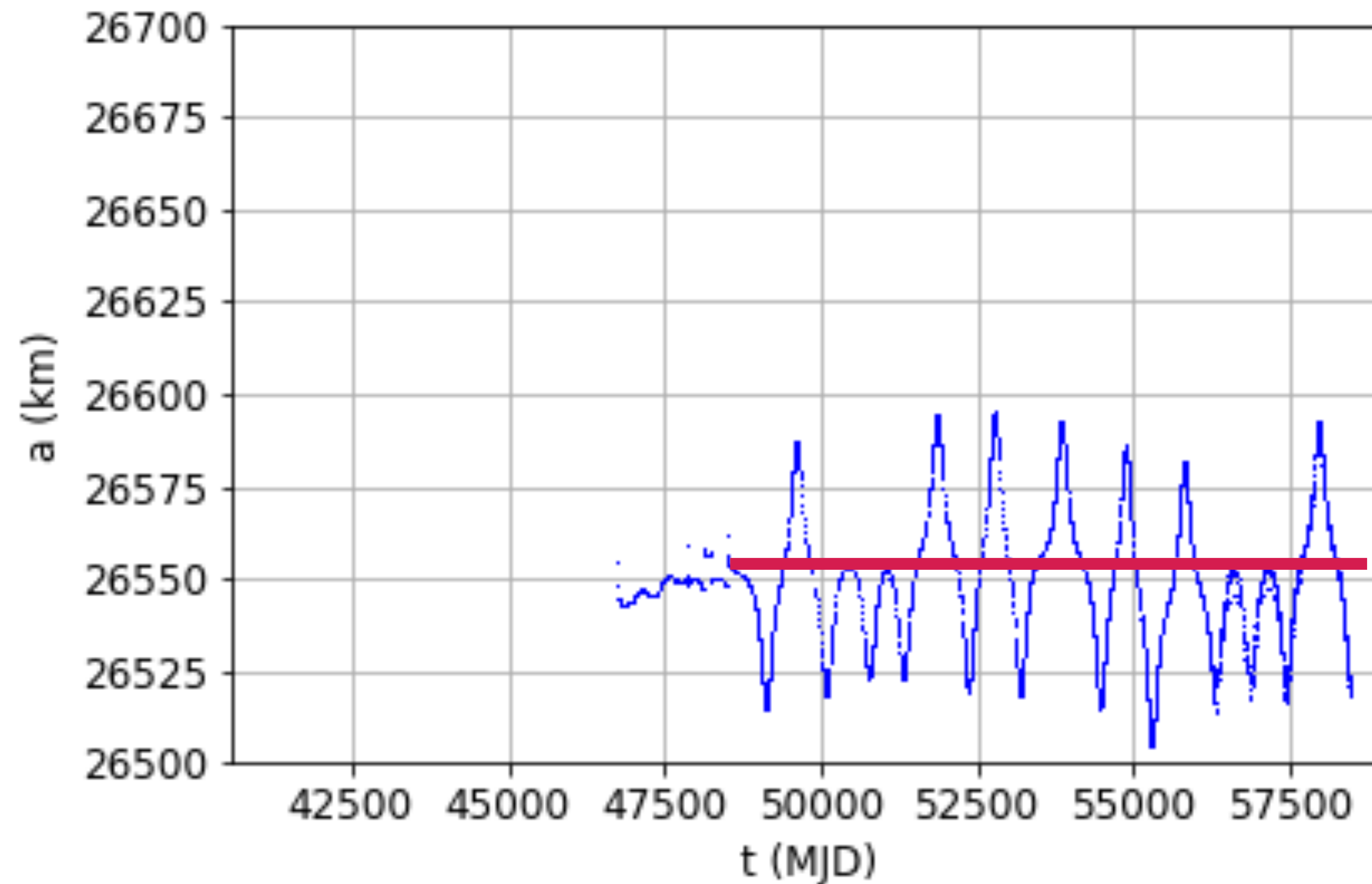
RESULTS

- **lunisolar perturbations must be considered for the evolution of ω and Ω ;**
- if the assumption of a constant is valid within a certain limit, it turns out that only some terms of the quadrupolar approximation is enough **to follow the evolution in eccentricity**. In particular, the **following components** are mandatory
 - $2\omega, 2\omega + \Omega, 2\omega - \Omega, \Omega$ (Sun and Moon)
 - $2\omega - \Omega_L, 2\omega + \Omega_L$
 - $2\omega + \Omega - \Omega_L, 2\omega - \Omega + \Omega_L$
 - $\Omega - \Omega_L$
- Lomb-Scargle analysis has detected the following **dominant frequencies**
 - $2\dot{\omega} + \dot{\Omega}$ (≈ 7 years)
 - $2\omega + \Omega_L$ (≈ 25 years)
 - $2\omega + \Omega - \Omega_L$ (≈ 11 years)

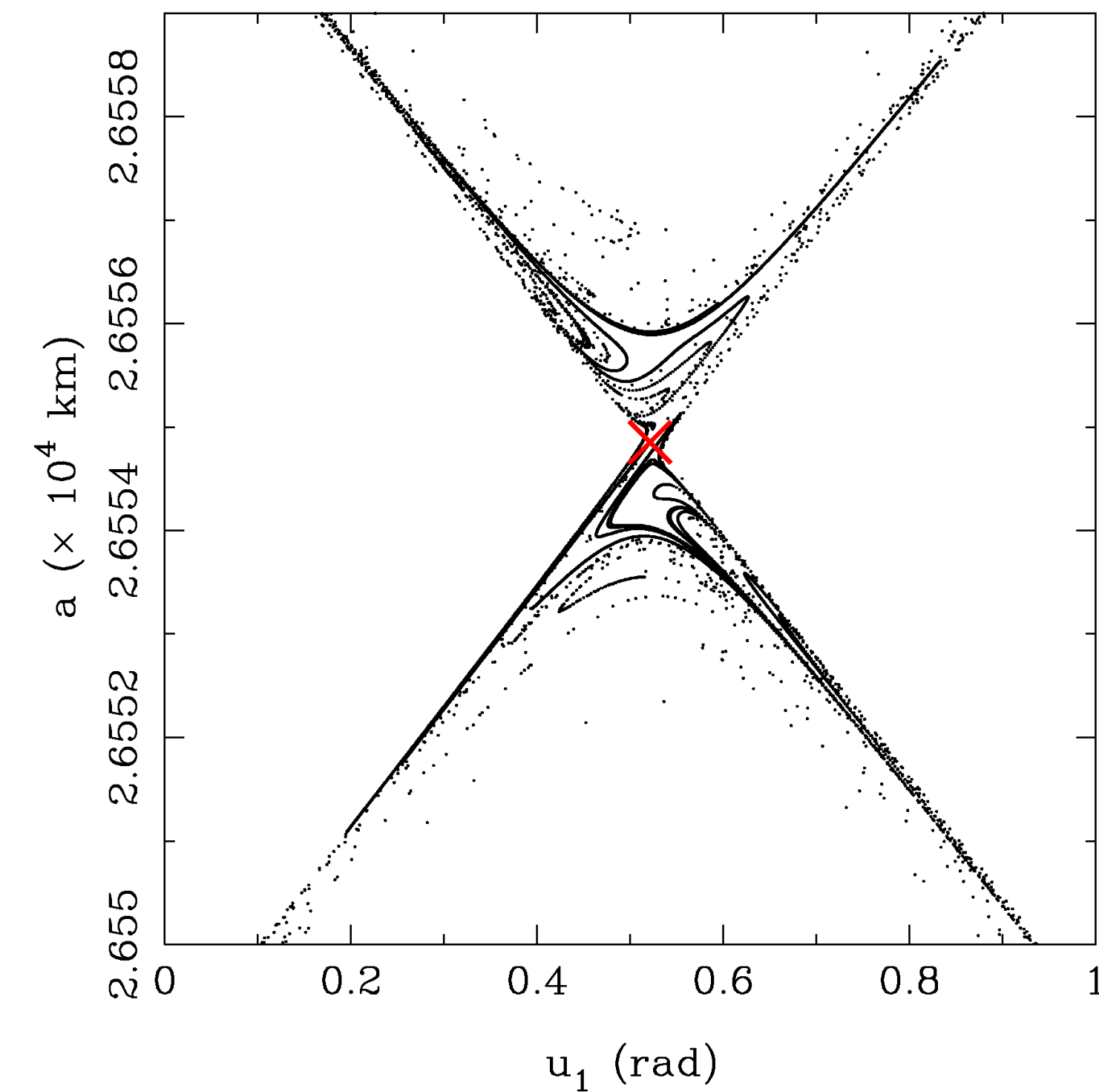
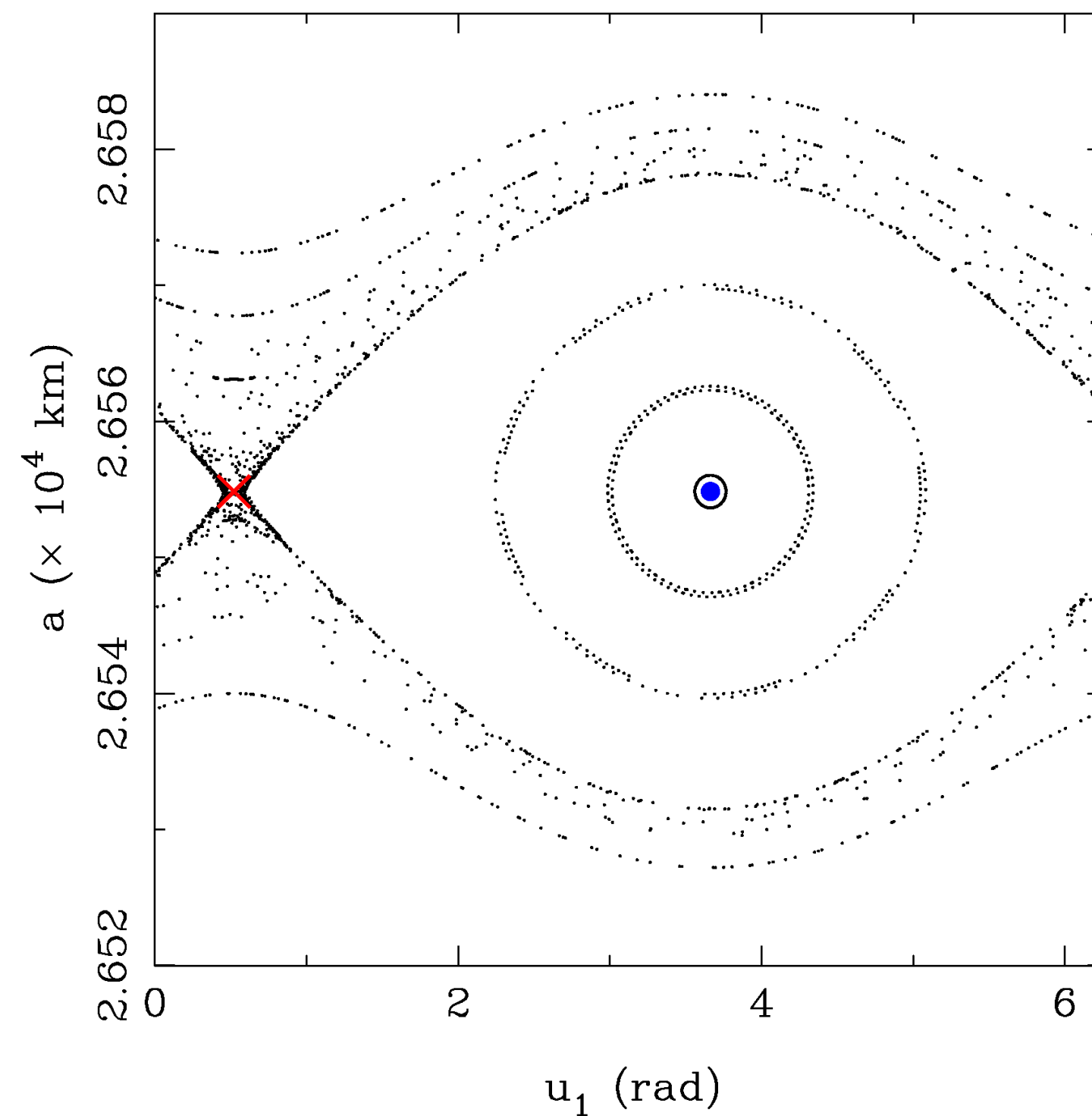
The dominant periodic terms can be explained looking to the ratio between the amplitude and the corresponding frequency.

SEMI-MAJOR AXIS EVOLUTION

SEMI-MAJOR AXIS INTERMITTENCY PHENOMENON



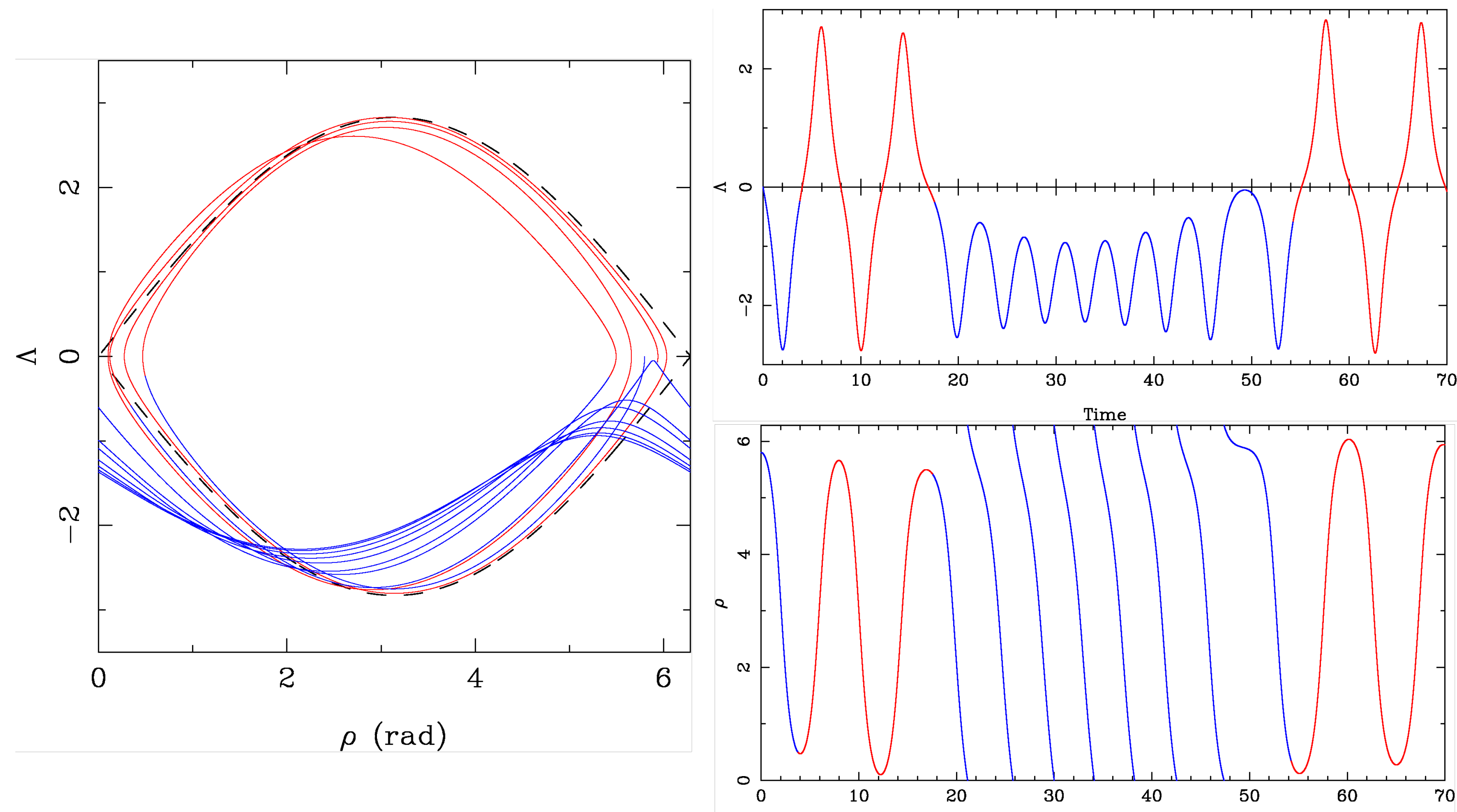
POINCARÉ SECTION



**Perturbative contributions to the Hamiltonian given by
the tesseral terms and the lunisolar perturbations**

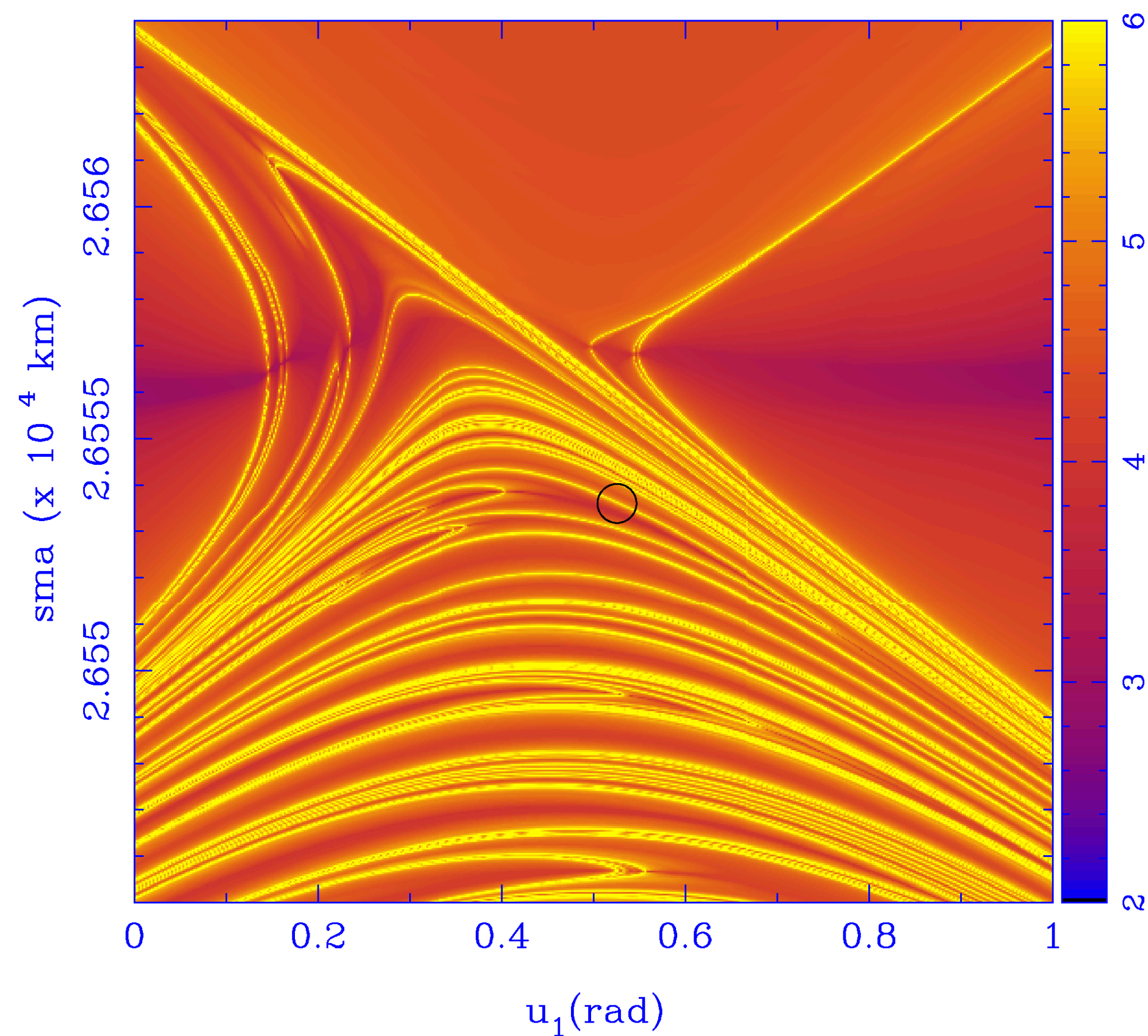
EXPLANATION OF THE INTERMITTENCY PHENOMENON

The chaotic nature close to the separatrices makes the satellite to jump from the libration to the circulation region

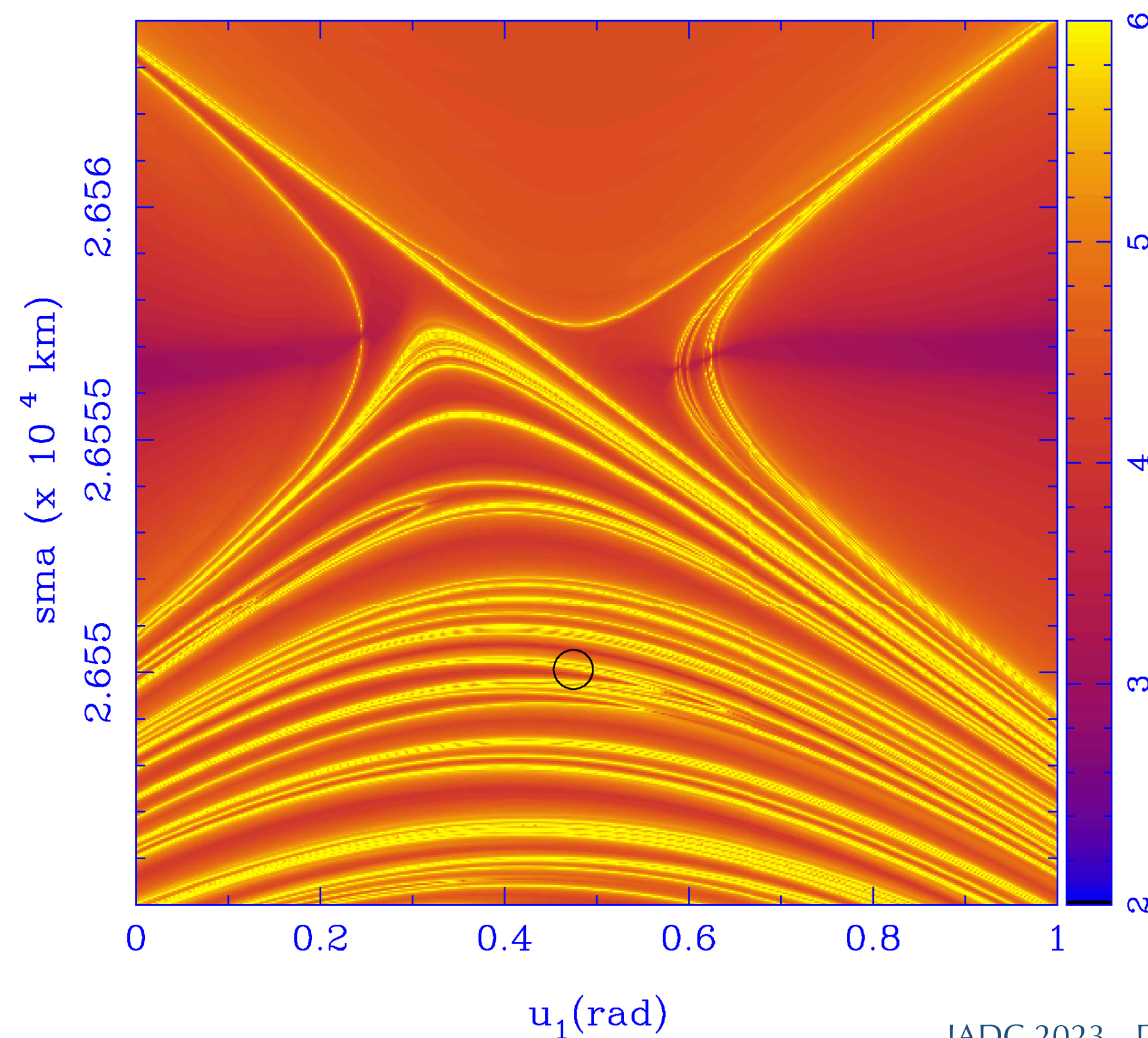


MOLNIYA 1-69 AND 1-87

Molniya M1-69

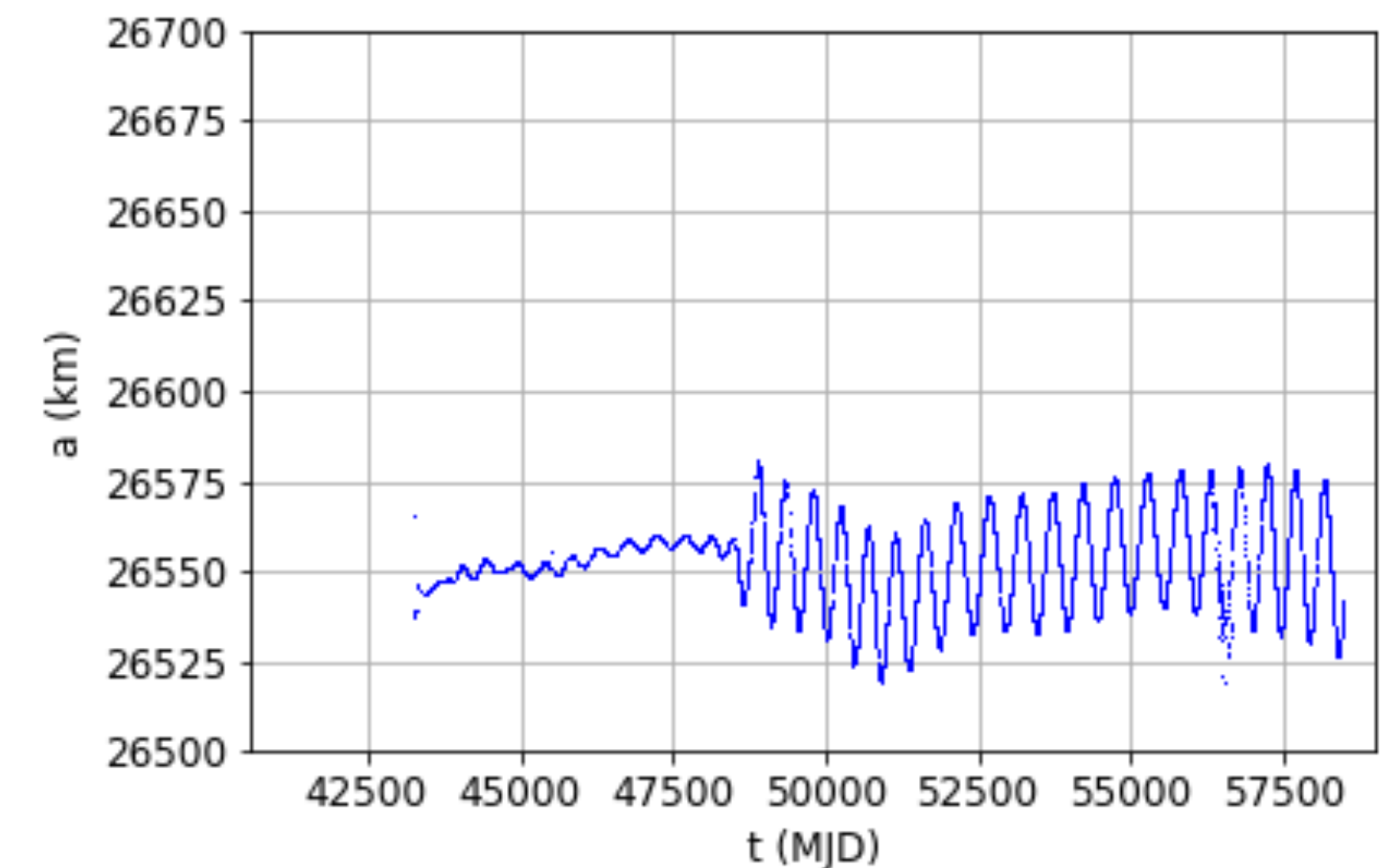


Molniya M1-87



OPEN POINTS - FURTHER DEVELOPMENTS

- change of amplitudes in the periodic behavior of the semi-major axis
- chaotic behavior in eccentricity?



- **lunisolar - tesseral - drag dynamical coupling** (also, any possible data from reentry campaigns?)
- **ω drift from the critical value?**
- **long-term evolution of debris fragments?**

PAST FRAGMENTATIONS?

TABLE 2.1 HISTORY OF SATELLITE BREAKUPS BY LAUNCH DATE (CONT'D)

Anz-Meador, P., Liou, J.C., History of On-orbit Satellite Fragmentations, 16th Edition, NASA/TP-20220019160, NASA Orbital Debris Program Office, December 2022.

SATELLITE NAME	INTERNATIONAL DESIGNATOR	US SATELLITE NUMBER	LAUNCH DATE	BREAKUP DATE	DEBRIS CATALOGED	DEBRIS LEFT	APOGEE (KM)	PERIGEE (KM)	INCLINATION (DEG)	ASSESSED CAUSE	ADDITIONAL INFORMATION
COSMOS 1278	1981-058A	12547	19-Jun-81	Dec-86	3	0	37690	2665	67.1	DELIBERATE	SELF-DESTRUCT
COSMOS 1285	1981-071A	12627	4-Aug-81	21-Nov-81	25	25	40100	720	63.1	DELIBERATE	SELF-DESTRUCT
COSMOS 1286	1981-072A	12631	4-Aug-81	29-Sep-82	2	0	325	300	65.0	UNKNOWN	COSMOS 699 CLASS
COSMOS 1305 R/B	1981-088F	12827	11-Sep-81	11-Sep-81	8	8	13795	605	62.8	PROPULSION	MOLNIYA FINAL STAGE
COSMOS 1306	1981-089A	12828	14-Sep-81	12-Jul-82	8	0	405	380	64.9	UNKNOWN	COSMOS 699 CLASS
COSMOS 1317	1981-108A	12933	31-Oct-81	25-28 Jan-84	11	11	39055	1315	62.8	DELIBERATE	SELF-DESTRUCT
METEOR 2-8	1982-025A	13113	25-Mar-82	29-May-99	53	53	960	935	82.5	UNKNOWN	
COSMOS 1348	1982-029A	13124	7-Apr-82	2-Sep-84	11	11	39200	1185	62.8	DELIBERATE	SELF-DESTRUCT
COSMOS 1355	1982-038A	13150	29-Apr-82	8-Aug-83	29	0	395	360	65.1	UNKNOWN	COSMOS 699 CLASS
COSMOS 1375	1982-055A	13259	6-Jun-82	21-Oct-85	62	59	1000	990	65.8	BATTERY	
COSMOS 1405	1982-088A	13508	4-Sep-82	20-Dec-83	32	0	340	310	65.0	UNKNOWN	COSMOS 699 CLASS
COSMOS 1408	1982-092A	13552	16-Sep-82	15-Nov-21	1760	990	490	465	82.6	COLLISION, DELIBERATE	Russian ASAT test
EKRAN 9	1982-093A	13554	16-Sep-82	23-Dec-83	1	1	35795	35788	0.7	BATTERY	
COSMOS 1423 R/B	1982-115E	13696	8-Dec-82	8-Dec-82	29	0	425	235	62.9	PROPULSION	MOLNIYA FINAL STAGE
ASTRON ULLAGE MOTOR	1983-020B	13902	23-Mar-83	3-Sep-84	1	0	1230	220	51.5	PROPULSION	PROTON-K BLOCK DM SOZ
NOAA 8	1983-022A	13923	28-Mar-83	30-Dec-85	7	1	830	805	98.6	BATTERY	
COSMOS 1456	1983-038A	14034	25-Apr-83	13-Aug-83	4	0	39630	730	63.3	DELIBERATE	SELF-DESTRUCT
COSMOS 1461	1983-044A	14064	7-May-83	11-Mar-85	189	3	890	570	65.0	UNKNOWN	COSMOS 699 CLASS

THANK YOU
