

Inter-Agency Space Debris Coordination Committee



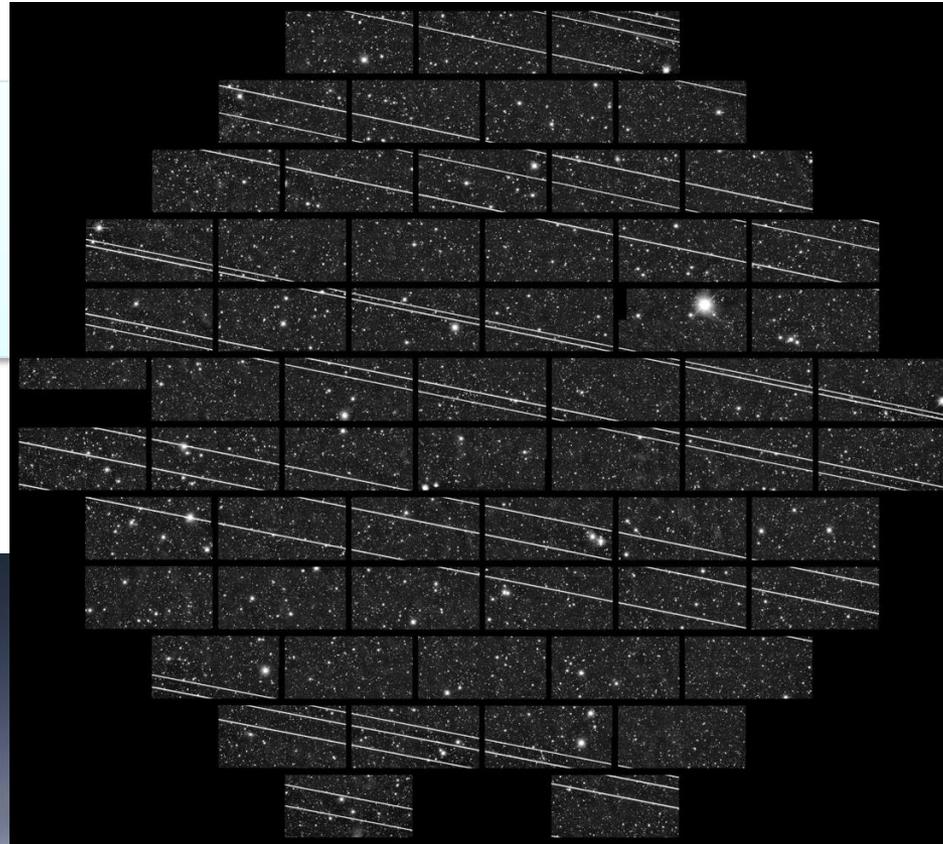
Mega-constellation photometry

Light pollution from starlink

2019-Nov-18 0800 UT

NSF Blanco 4.0-m
telescope

Cerro Tololo, Chile

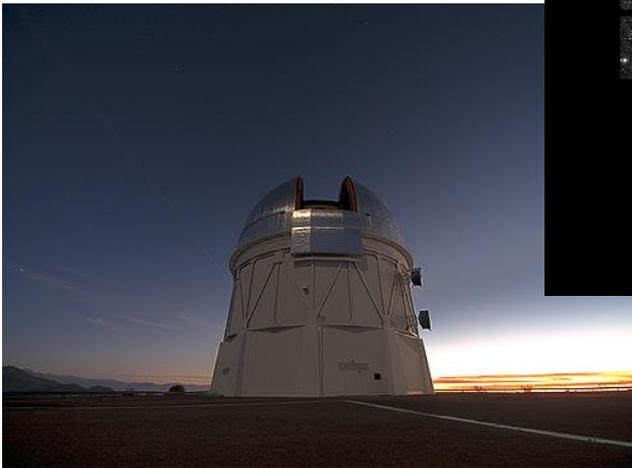


19 Starlinks

~4 sec

4 x moon
diameter

Starlinks
observed
not
intentionally



**Courtesy: Prof. Patrick
Seitzer, UMICH**

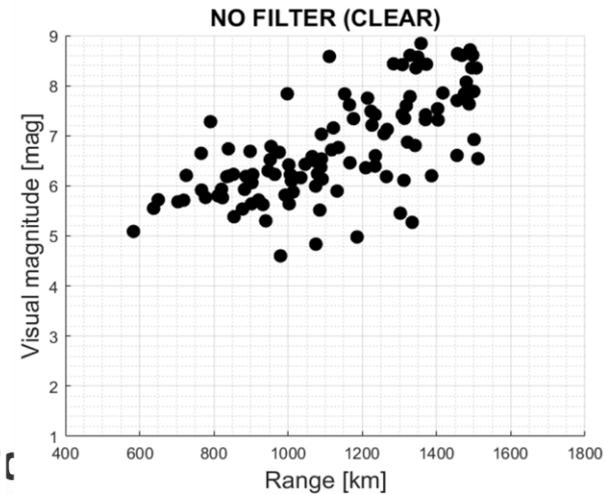
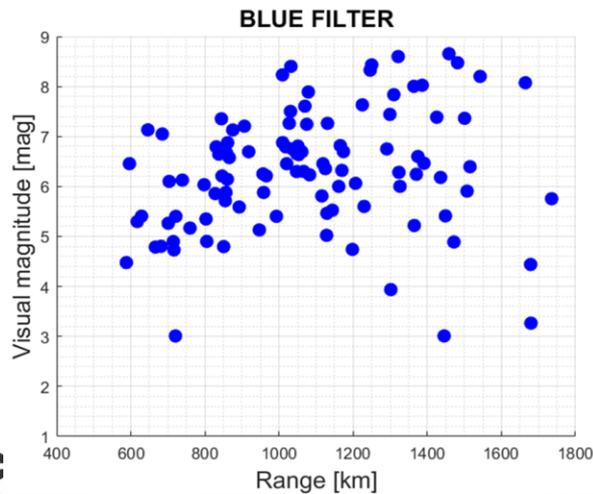
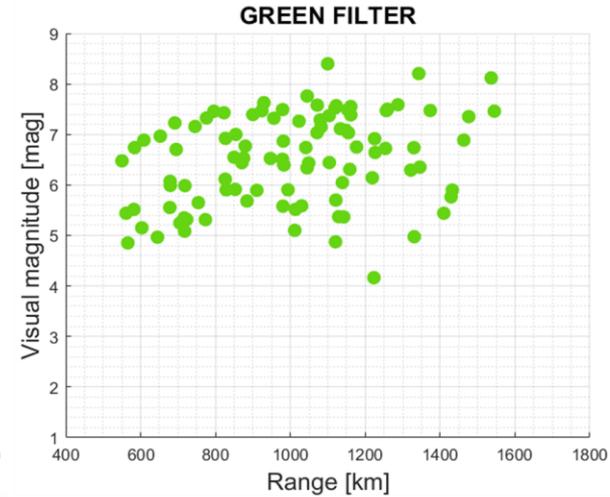
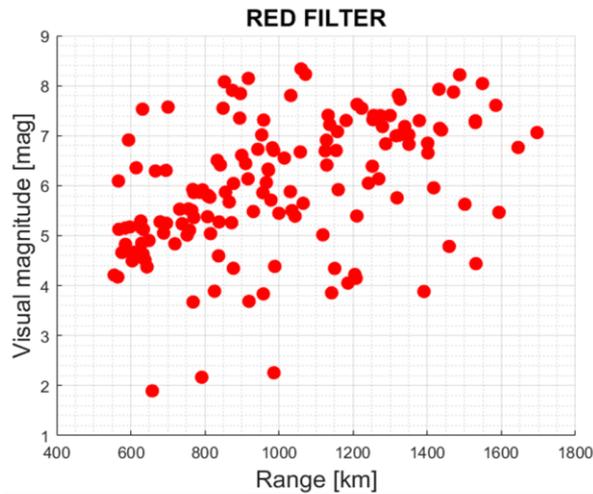
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02/05/2023

Starlink observations

Range vs. magnitude

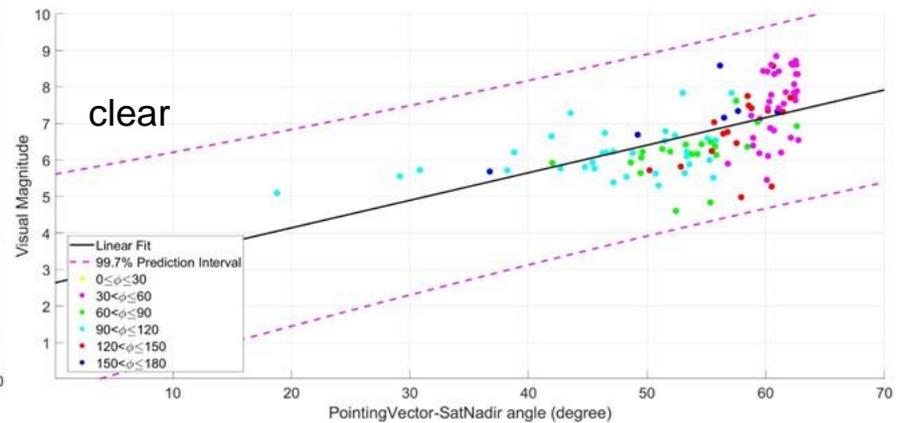
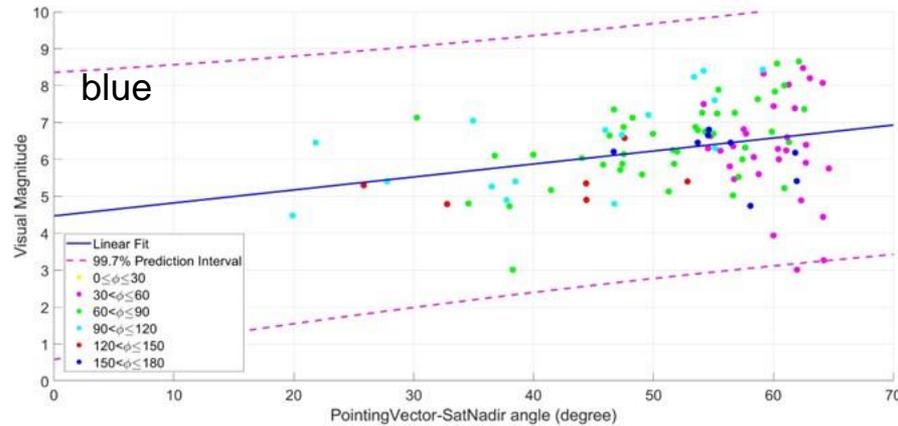
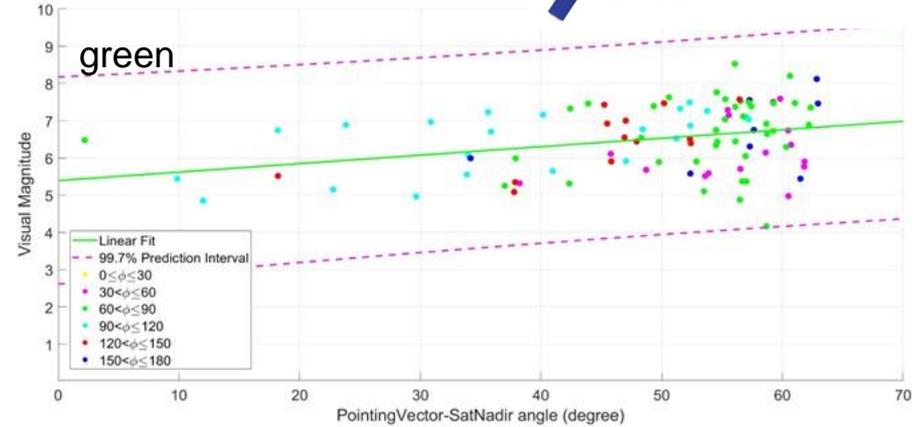
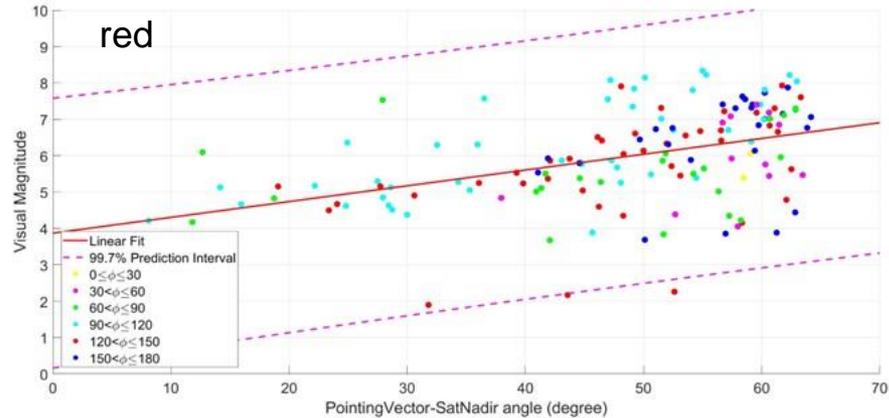
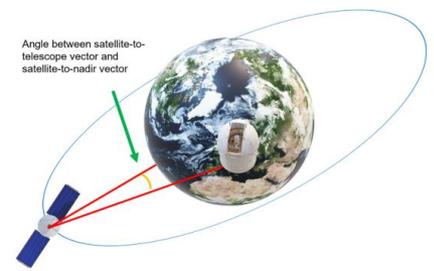


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Starlink observations

Magnitude vs phase angle



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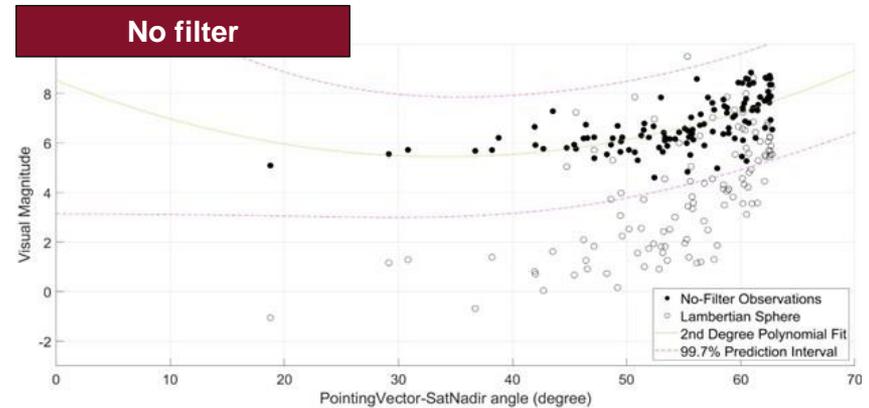
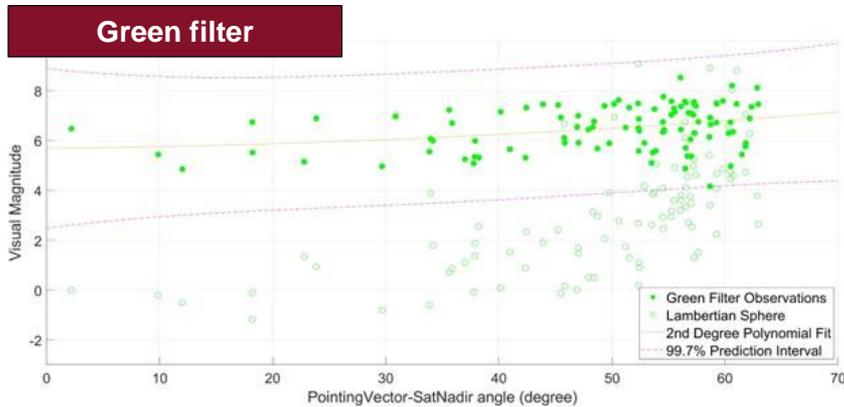
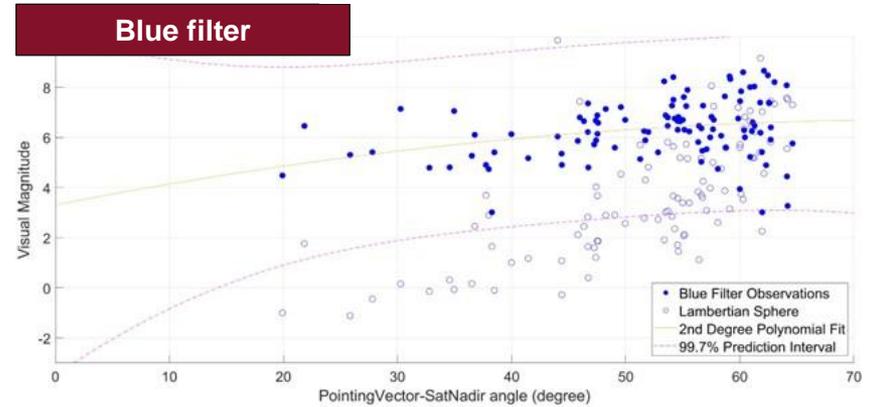
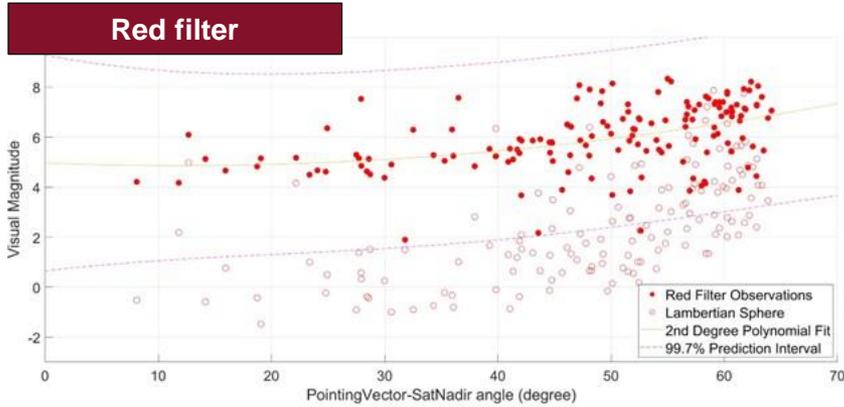
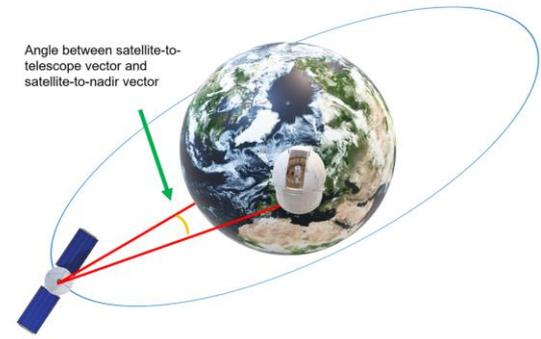
Photometric characterization of Starlink satellite tracklets using RGB filters
 Shariar Hadji Hossein^{a,*}, Lorenzo Cimino^a, Matteo Rossetti^a, Gaetano Zarcone^a, Lorenzo Mariani^a, Federico Curiano^a, Mascia Bucciarelli^a, Patrick Seitzer^a, Fabio Santoni^a, Alessandra Di Cecco^b, Fabrizio Piergentili^a

^a Department of Mechanical and Aerospace Engineering (DIMEAS), University of Rome "La Sapienza", Via Salaria, 16, 00186 Rome (RM), Italy
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Starlink observations

Magnitude vs Nadir Angle, different phase angles

Starlinks result less bright than Lambertian spheres



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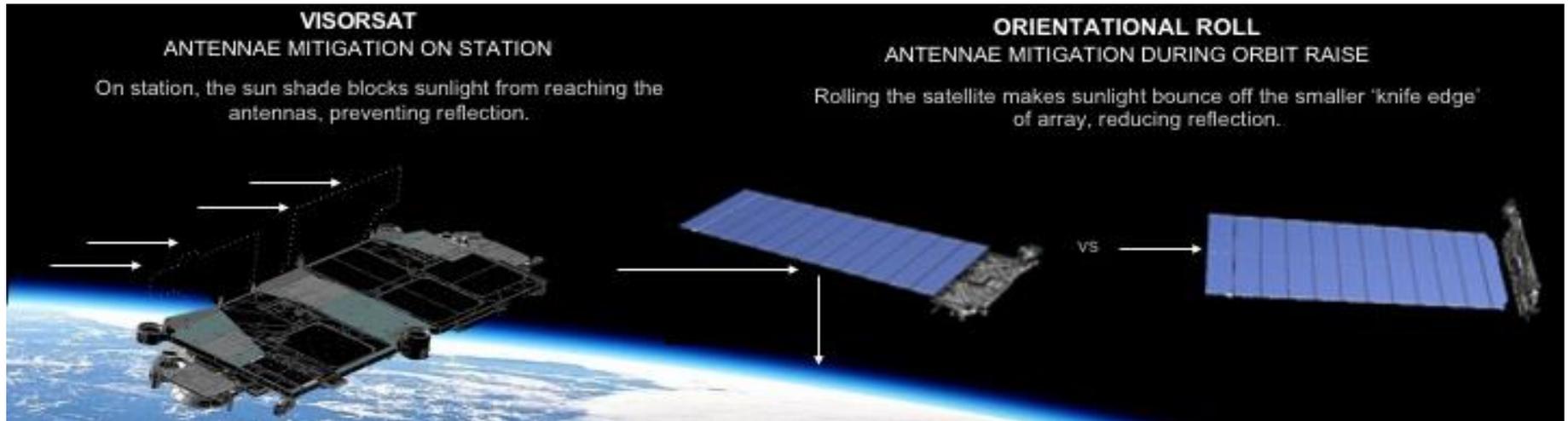
Starlink observations visorsat

Observation system and overview of the collected data



Lens diameter	250 mm
Focal length	1200 mm
Field of View	1.73° x 1.15°
Optical sensor	CCD
Mount type	Equatorial

Filter	N frames	VisorSat	Mean Magnitude
Red	148	25	8.2350
Green	98	33	8.1607
Blue	103	29	8.3335
No filter	111	24	8.2128
		Non-VisorSat	Mean magnitude
		123	7.0129
		65	7.2251
		74	7.0519
		87	7.2259

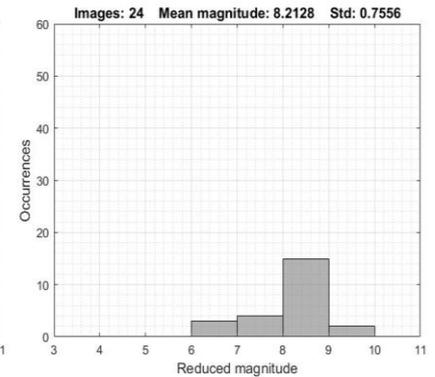
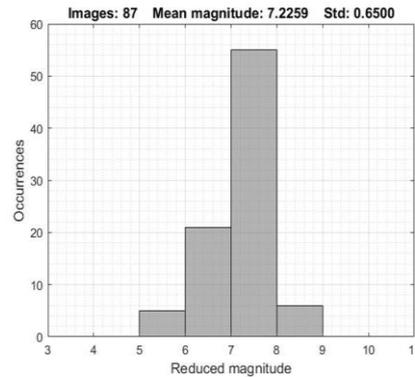
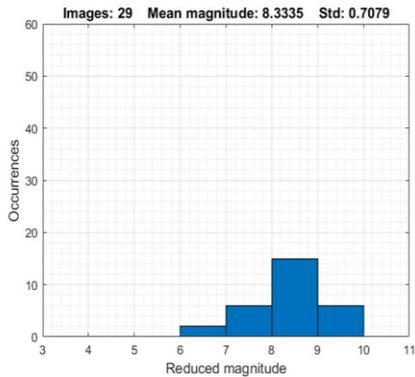
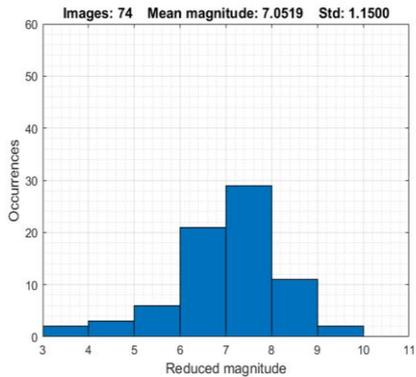
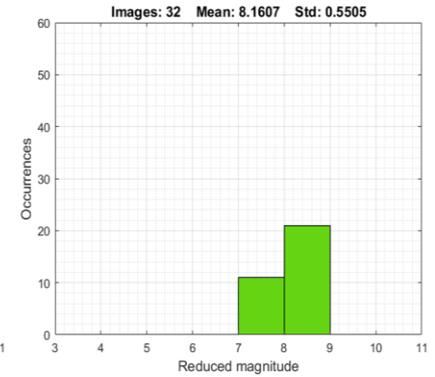
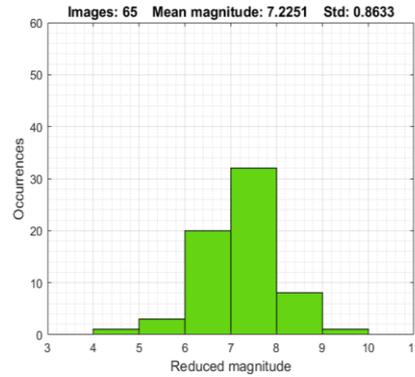
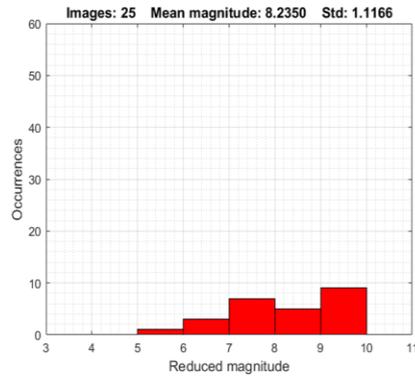
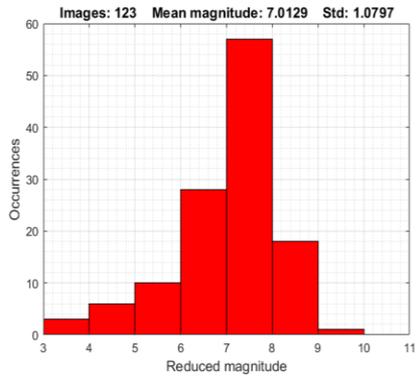


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Starlink photometric observations

Comparison between magnitudes of VisorSat and non-VisorSat Starlinks



23/09/2021

4

73rd International Astronautical Congress (IAC), Paris, France, 18-22 September 2022.
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IAC-22-D1,IPB,2,x73606

Photometric analysis for testing Starlink solutions to light reflection mitigation

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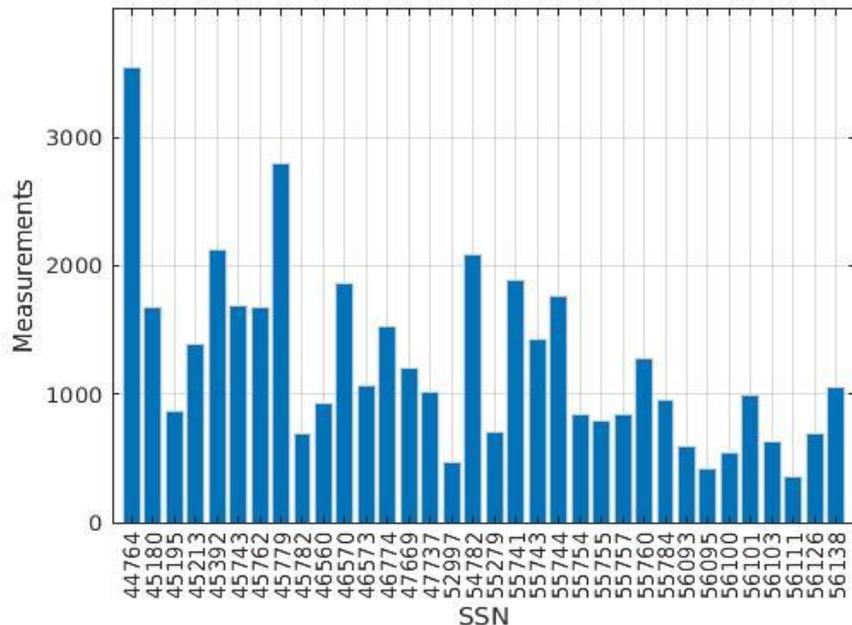
Lorenzo Cimino**, Shariar Hadji Hossein*, Lorenzo Mariani*, Gaetano Zarcone*, Matteo Rossetti*,
Mascia Bucciarelli*, Fabio Santoni*, Alessandra Di Cecco*, Patrick Seitzer*, Fabrizio Piergentili*



Starlink observations orbital assessment

TLE tracking

- in orbit raising and satellites in atmospheric re-entry
- Study on the impact of the mega-constellations on the night sky



2023 BOXSCORE	
Starlink observed	34
Number of observations	51
Descending phase	23
Ascending phase	28
Measurements	25000+
Filter	V-Johnson (peak 540 nm)

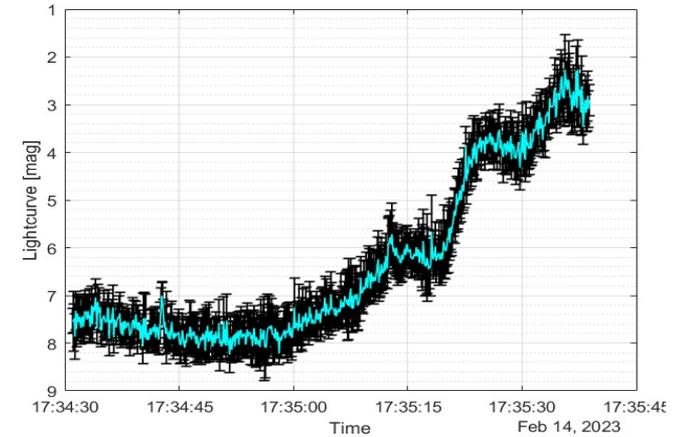
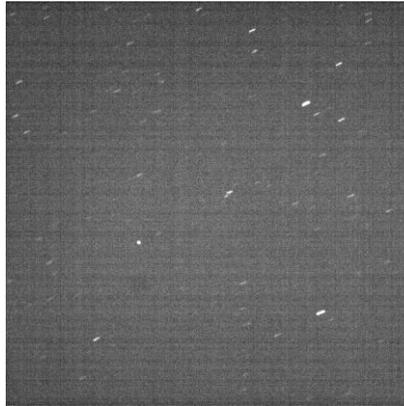
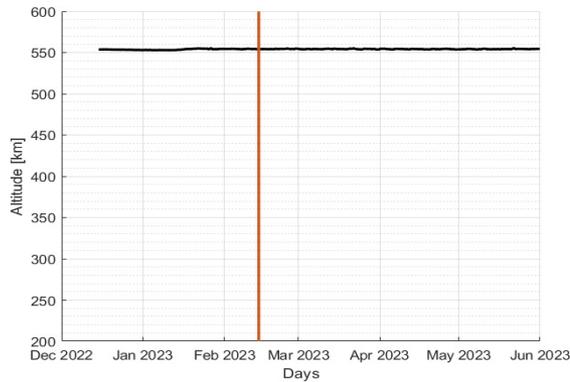
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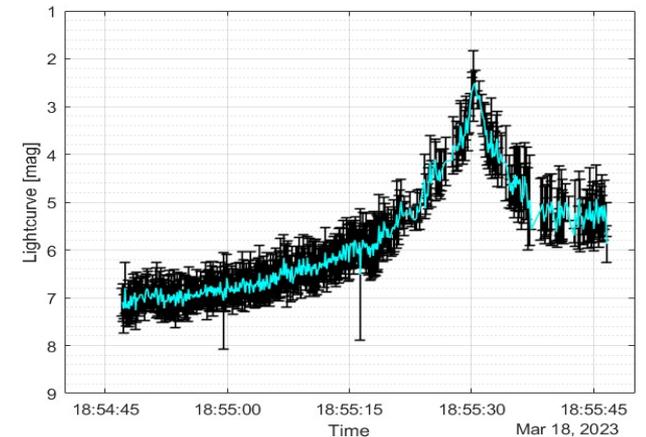
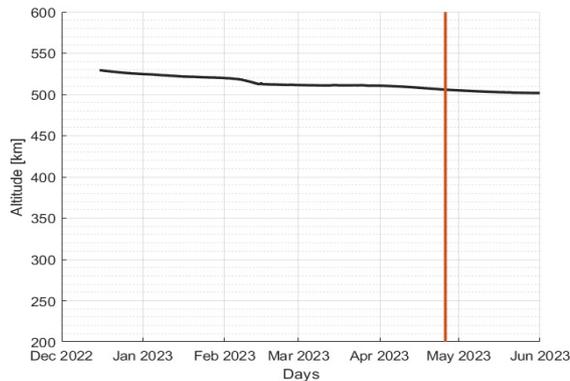
Starlink observations orbital assessment

Constant/Descending phase in TLE tracking

46674 – 2023/02/14



46570 – 2023/03/18



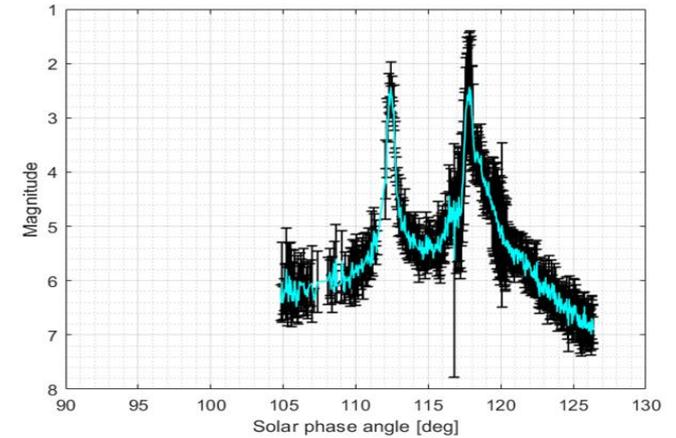
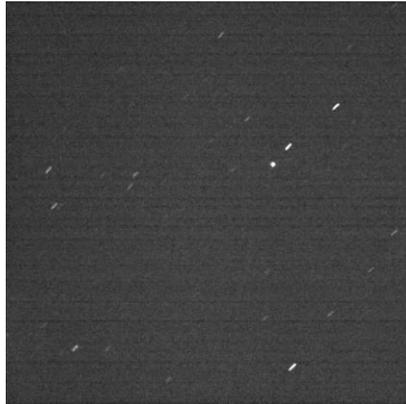
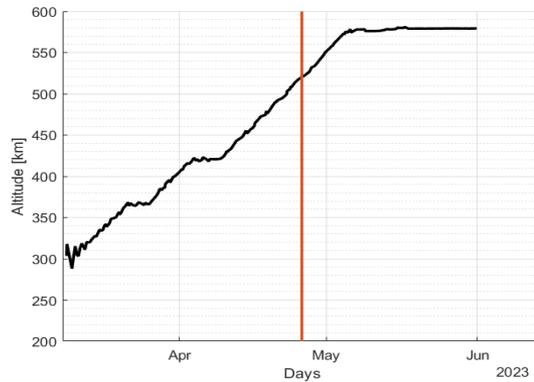
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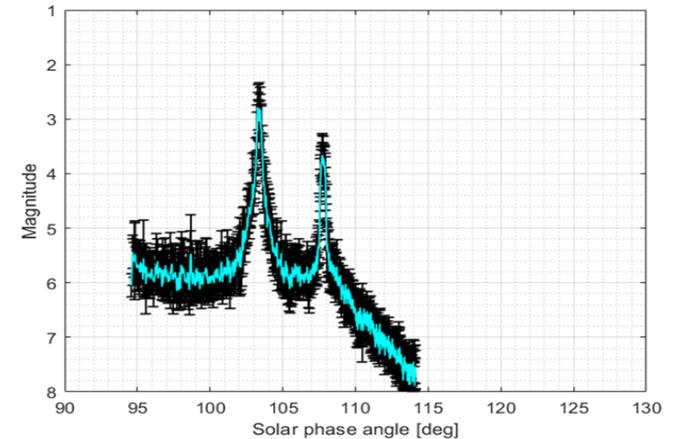
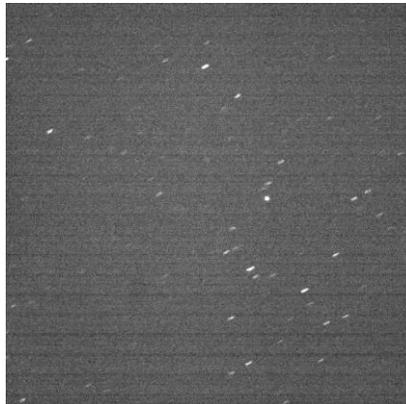
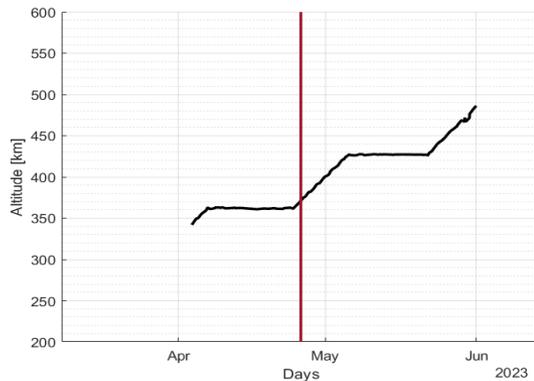
Starlink observations orbital assessment

Raising phase in TLE tracking

55760 – 2023/04/26



56138 – 2023/04/26



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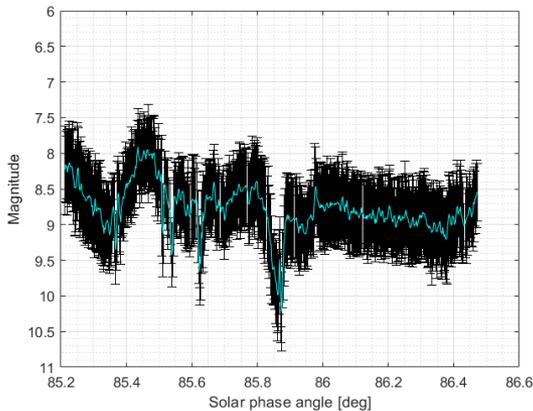


Oneweb observations

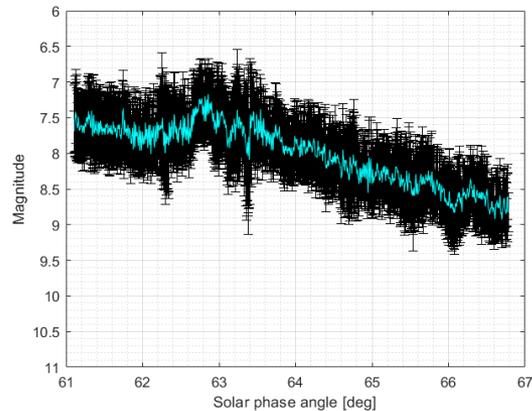
phase angle vs. magnitude

Collaboration with University of Michigan

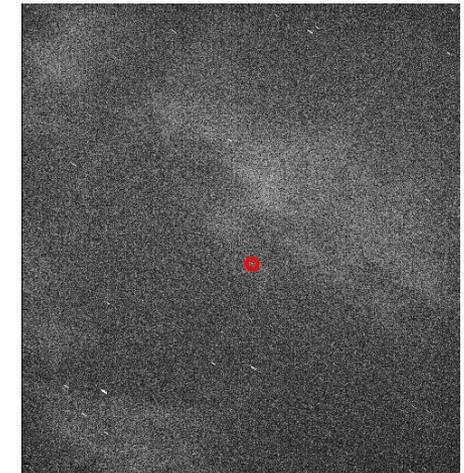
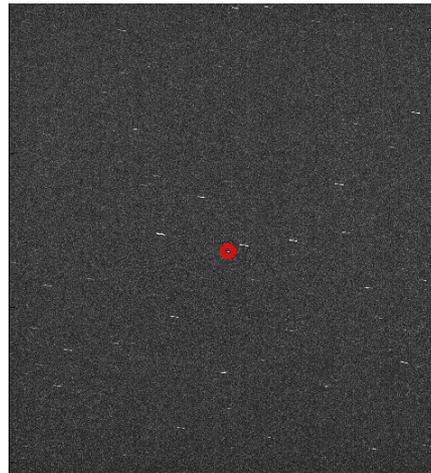
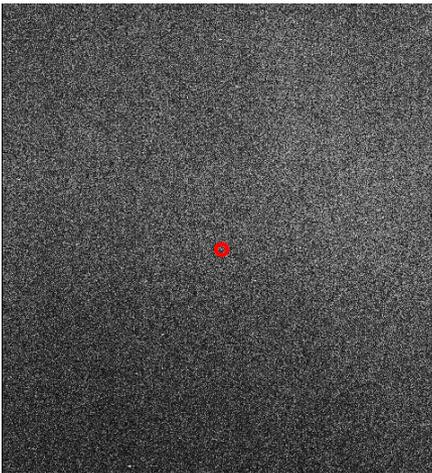
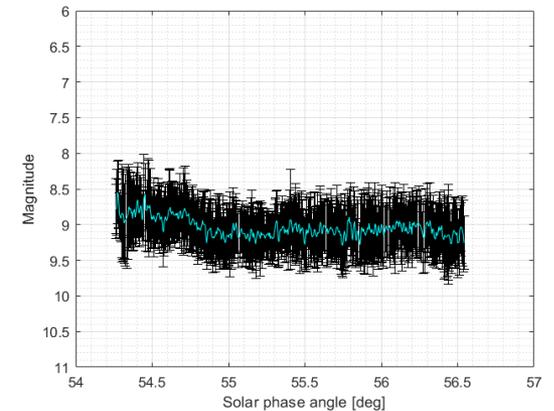
47265 – 2022/08/13



45445 – 2022/08/13



45143 – 2022/08/13

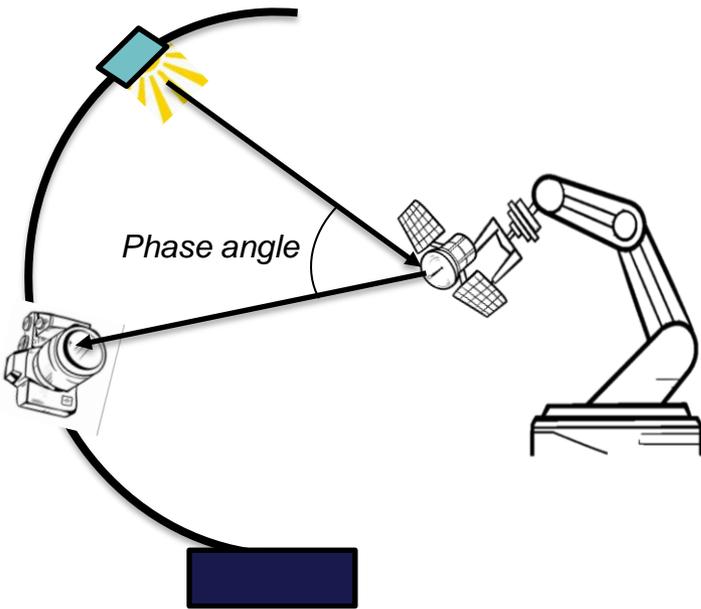


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Involve laboratory facilities

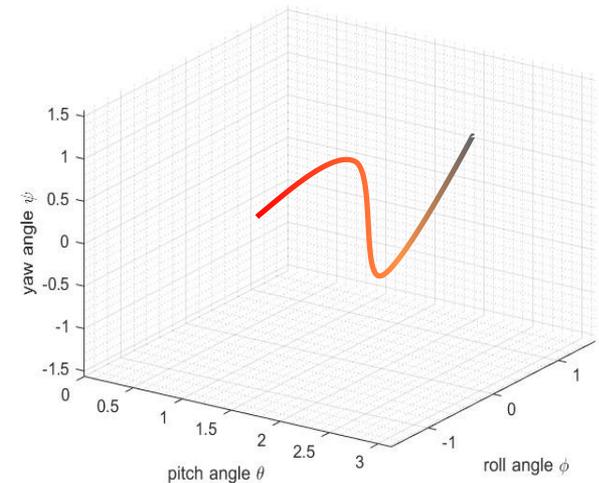
ARM4NDO



Satellite attitude variations during a simulated scenario using the robotic arm

Generation of a light curve wrt satellite attitude:
 $(\phi, \theta, \psi) \rightarrow \text{brightness}$

Light curve inversion:
processing simulations data for associating brightness values to attitude initial condition



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Involve laboratory facilities

ARM4NDO

Arm for Near Distance Observations

1. Attitude reconstruction from light curve

Satellite attitude variations during a real passage simulated using the robotic arm

2. Support the design of low impact satellites for the reduction of light pollution

The realization of a reflection model before launch permits to implement mitigation actions in order to prevent excessive brightness for constellations such as IRIDE, Starlink and OneWeb

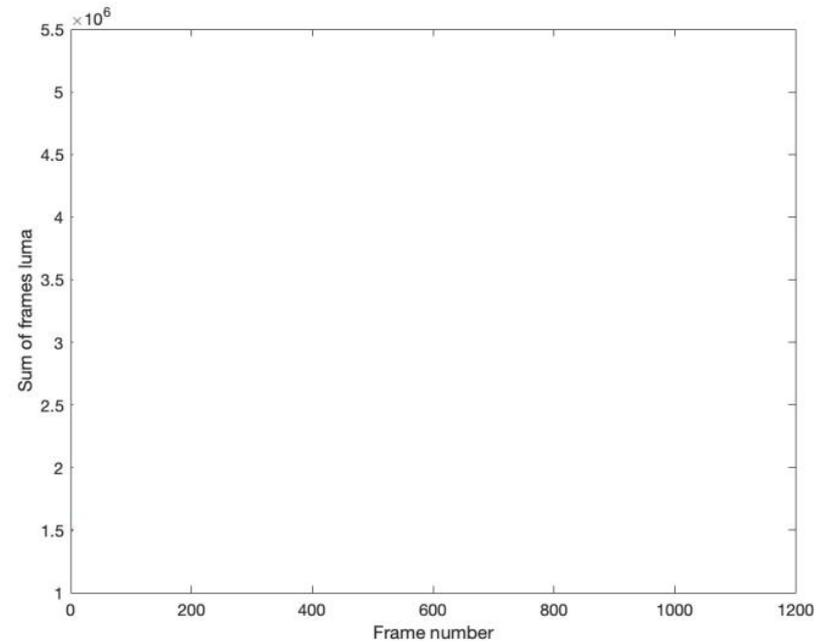
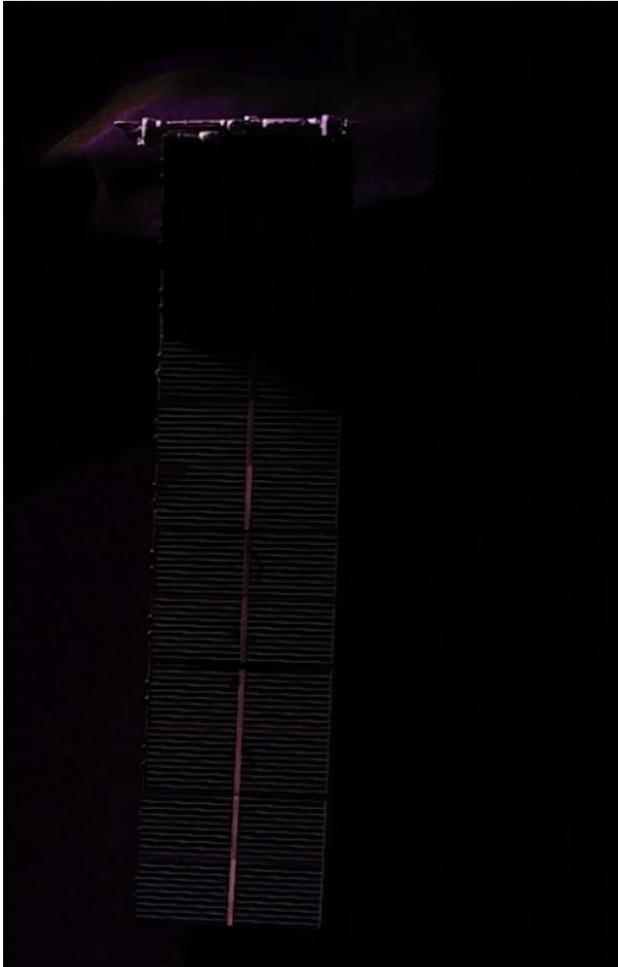


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Involve laboratory facilities

Test

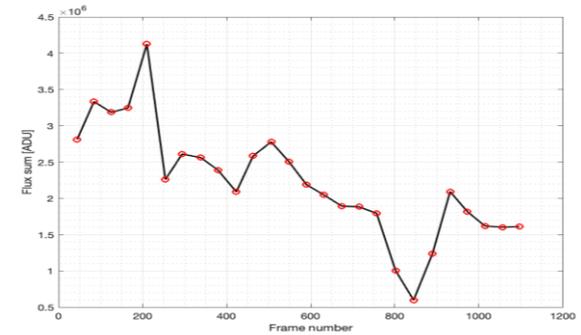
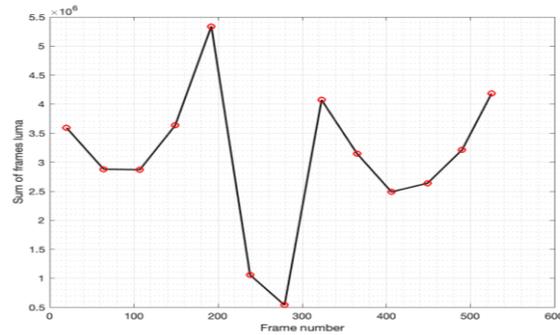
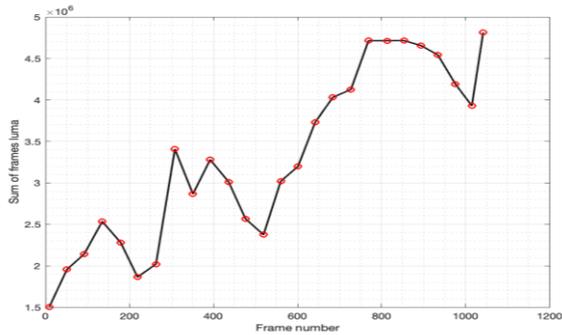
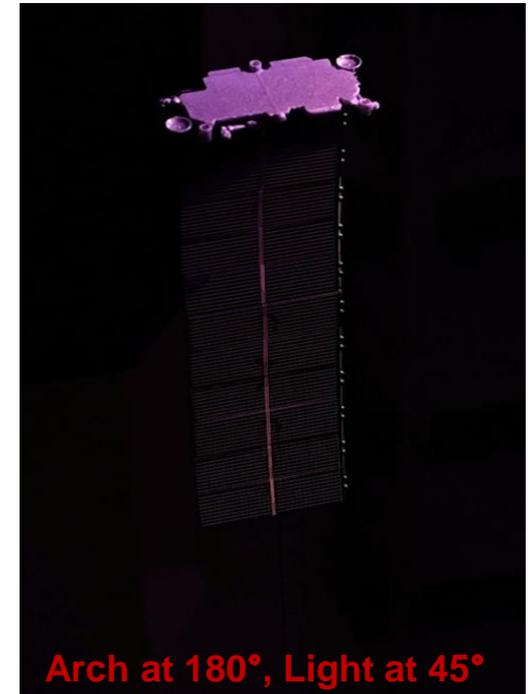
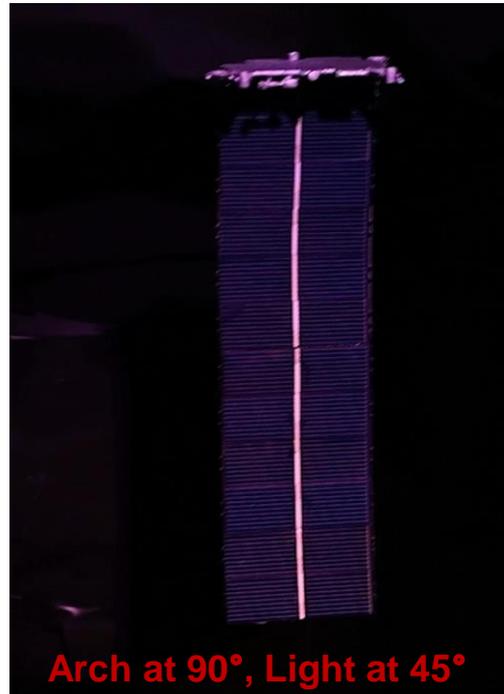
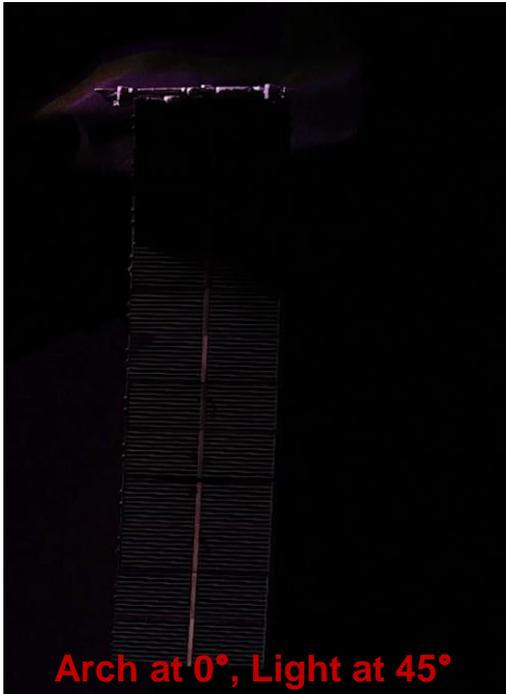


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Involve laboratory facilities

Test



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Conclusion

- Mega-constellations poses a serious **space sustainability** risk
- New **standards and guidelines** are being studied from different institution
- Paramount importance of satellites in mega-constellation for:
 - Evaluating effectiveness of **mitigation systems** (paints, visors, maneuver...)
 - Evaluating **polluting effects** on different bands (optical, infrared, microwaves...)
 - Evaluate the effects of different **observation geometries**, distance and phase angle
 - Evaluate **new design solutions** to mitigate polluting effects through ground characterization (lab measurements)

- **Measurements of Mega constellations target are easy to obtain**

- **Extracting reliable informatons for Megaconstellations is not trivial –strong effect from geometry, satellite design updates and orbital phase**

- **Different constellation could have different behaviour**

- **Using lab facilities to produce experimental light curve could be useful - digital twin seems not to be enough accurate or needs to be validated from experimental data**

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